

1
2
3
4
5
6
7
8
9
10
11

Impacts of Heterogeneous Chemistry on Vertical Profiles of Martian Ozone

M. A. J. Brown¹, M. R. Patel^{1,2}, S. R. Lewis¹, J. A. Holmes¹, G. J. Sellers¹,
P. M. Streeter¹, A. Bennaceur¹, G. Liuzzi^{3,4}, G. L. Villanueva³, A. C.
Vandaele⁵

¹The Open University, Milton Keynes, U.K.

²Space Science and Technology Department, Science and Technology Facilities Council, Rutherford
Appleton Laboratory, Oxfordshire, U.K.

³Planetary Systems Laboratory, NASA Goddard Space Flight Center, Greenbelt, MD, USA

⁴Department of Physics, American University, Washington, DC, USA

⁵Royal Belgian Institute for Space Aeronomy (BIRA-IASB), Brussels, Belgium

12
13
14
15
16
17
18

Key Points:

- Ozone and water ice profiles show a positive vertical correlation, contrary to global ozone and water vapour total column anti-correlation
- Heterogeneous chemistry increases ozone abundance at altitudes where water ice is present, matching locations where ozone is underpredicted
- High hydroxyl radical abundance, assumed proportional to high water vapour abundance, masks the effects of heterogeneous chemistry on ozone

Corresponding author: M.A.J. Brown, megan.brown@open.ac.uk

19 **Abstract**

20 We show a positive vertical correlation between ozone and water ice using a ver-
 21 tical cross-correlation analysis with observations from the ExoMars Trace Gas Orbiter’s
 22 NOMAD instrument. We find this is particularly apparent during the first half of Mars
 23 Year 35 ($L_S = 0^\circ - 180^\circ$) at high southern latitudes, when the water vapour abundance
 24 is low. This contradicts the current understanding that ozone and water are, in general,
 25 anti-correlated. However, our simulations with gas-phase-only chemistry using a 1-D model
 26 show that ozone concentration is not influenced by water ice. Heterogeneous chemistry
 27 has been proposed as a mechanism to explain the underprediction of ozone in global cli-
 28 mate models (GCMs) through the removal of HO_x . We find improving the heterogeneous
 29 chemical scheme causes ozone abundance to increase when water ice is present, better
 30 matching observed trends. When water vapour abundance is high, there is no consistent
 31 vertical correlation between observed ozone and water ice and, in simulated scenarios,
 32 the heterogeneous chemistry does not have a large influence on ozone. HO_x , which are
 33 by-products of water vapour, dominate ozone abundance and mask the effects of het-
 34 erogeneous chemistry on ozone. This is consistent with gas-phase-only modelled ozone,
 35 showing good agreement with observations when water vapour is abundant. High wa-
 36 ter vapour abundance masks the effect of heterogeneous reactions on ozone abundance
 37 and makes adsorption of HO_x have a negligible impact on ozone. Overall, the inclusion
 38 of heterogeneous chemistry improves the ozone vertical structure in regions of low wa-
 39 ter vapour abundance, which may partially explain GCM ozone deficits.

40 **Plain Language Summary**

41 Ozone gas is found in small quantities in the martian atmosphere, highly variable
 42 both in time and space. Ozone quantity is controlled by photochemical reactions involv-
 43 ing other species too difficult to detect with remote sensing. Two main ways ozone is de-
 44 stroyed in the martian atmosphere are by: 1. absorbing ultraviolet sunlight; 2. react-
 45 ing with hydroxyl radicals, highly reactive chemical formed by water vapour absorbing
 46 ultraviolet sunlight. The latter leads to a well-known anti-correlation between ozone and
 47 water vapour. Ozone is underpredicted in climate models, implying chemical reactions
 48 are missing or incorrect in models. We investigate reactions between hydroxyl radicals
 49 and water ice as a potential explanation for the ozone underprediction by using a model
 50 and observed vertical profiles. We find observed ozone and water ice have a positive ver-
 51 tical correlation and, when simulating a model with improved chemistry, ozone abun-
 52 dance increases at altitudes where water ice is present due to the decrease in hydroxyl
 53 radicals. However, in areas where water vapour is abundant, no such correlation is seen
 54 and the chemistry has little impact on modelled ozone. This is due to large abundances
 55 of hydroxyl radicals which inhibit the increase in ozone caused by the addition of the im-
 56 proved chemistry.

57 **1 Introduction**58 **1.1 Background**

59 Ozone on Mars was first detected in 1969 by Mariner 7 and later by Mariner 9. It
 60 was found to vary seasonally in both hemispheres, with a greater abundance in local win-
 61 ter which decreased during the local summer (Barth et al., 1973). Other instruments,
 62 such as SPICAM (Spectroscopy for the Investigation of the Characteristics of the At-
 63 mosphere of Mars) aboard Mars Express (Bertaux et al., 2006), MARCI (Mars Colour
 64 Imager) aboard Mars Reconnaissance Orbiter (Bell III et al., 2009), and NOMAD (Nadir
 65 and Occultation for Mars Discovery) aboard the ExoMars Trace Gas Orbiter (TGO) (M. R. Pa-
 66 tel et al., 2017; Vandaele et al., 2018) later confirmed the seasonal variability, and ob-

67 servations show that the highest abundance of ozone occurs at high latitudes (e.g. Perrier
68 et al. (2006); Clancy et al. (2016); M. R. Patel et al. (2021); Khayat et al. (2021)).

69 In the martian atmosphere, ozone is a trace gas highly sensitive to direct (photol-
70 ysis) and indirect (reactions with photolysed products) photochemical reactions. Due
71 to its sensitivity to other chemical species and its relatively short lifetime (2-3 hours on
72 the dayside), ozone is often a useful indicator of the chemical reactions occurring in the
73 atmosphere (Clancy & Nair, 1996). One of the main destructive pathways of ozone is
74 via reactions with hydroxyl radicals (OH, HO₂; HO_x) and atomic hydrogen, H, which
75 are a by-product of the photolysis of water vapour (Shimazaki & Shimizu, 1979; Clancy
76 & Nair, 1996; Lefèvre et al., 2004). HO_x are highly reactive and cause a set of chain re-
77 actions which produce more HO_x and lead to further ozone destruction. Their high re-
78 activity makes them a key component in understanding the stability of the martian at-
79 mosphere as they catalyse the recombination of atomic oxygen and carbon monoxide to
80 form carbon dioxide (McElroy & Donahue, 1972; Clancy & Nair, 1996). The destruc-
81 tive pathway of ozone caused by HO_x results in a photochemical anti-correlation between
82 ozone and water vapour. As a result, the seasonal variation in ozone is influenced by the
83 fluctuation in water vapour throughout the year.

84 The seasonal cycle and asymmetric distribution of water vapour between the two
85 hemispheres leads to a larger ozone abundance at high southern latitudes during south-
86 ern winter, compared to high northern latitudes during northern winter (Perrier et al.,
87 2006; Montmessin & Lefèvre, 2013). So far, the highest recorded abundances of ozone
88 have occurred at high southern latitudes (> 55° S) between the northern spring equinox
89 and the northern autumnal equinox ($L_S = 0^\circ - 180^\circ$). Global climate models (GCMs)
90 are also in agreement with these findings (Lefèvre et al., 2004; Perrier et al., 2006; Clancy
91 et al., 2016; J. A. Holmes et al., 2018).

92 Similarly, water ice has been detected at high latitudes between $L_S = 0^\circ - 180^\circ$
93 (e.g. Smith (2004); Benson et al. (2010, 2011); Wolff et al. (2019)). During the north-
94 ern and southern winters, water vapour condenses over the polar regions and forms water-
95 ice clouds known as the North and South Polar Hood respectively (Benson et al., 2010,
96 2011). The North Polar Hood lasts from $L_S = 150^\circ$ to $L_S = 30^\circ$ the following year, while
97 the South Polar Hood is observed between $L_S = 10^\circ - 200^\circ$. Clancy et al. (2016) and
98 Daerden et al. (2019) showed that the greatest underprediction of ozone occurs between
99 $L_S = 0^\circ - 180^\circ$ at high (> 60° N/S) latitudes, as compared against MARCI observa-
100 tions. Lefèvre et al. (2008) found that ozone was underpredicted throughout the year
101 across all latitudes, although the largest differences between the modelled and observed
102 data from SPICAM were also between $L_S = 0^\circ - 180^\circ$ at high northern and southern
103 latitudes. M. R. Patel et al. (2021) showed that the greatest underprediction of ozone
104 occurs in the southern hemisphere at low altitudes in the vertical profile also between
105 $L_S = 0^\circ - 180^\circ$.

106 Due to the chemical sensitivity of ozone, an ozone deficit can be used as a proxy
107 for missing or undeveloped chemical reactions in atmospheric models (Nair et al., 1994;
108 Lefèvre et al., 2008). Anbar et al. (1993) suggested ozone abundance could be increased
109 by adding heterogeneous chemistry, in the form of incorporating the adsorption of HO_x
110 onto the surface of water ice. They used a 1-dimensional (1-D) model to simulate this
111 chemistry, and found that the water ice acted as a sink for HO_x, reducing the destruc-
112 tion rate of ozone and resulting in an enhanced ozone abundance. Lefèvre et al. (2004)
113 briefly mentioned that heterogeneous chemistry could increase the ozone abundance as
114 it was still underpredicted in a 3-D model. This was further developed by Lefèvre et al.
115 (2008) who included 2 heterogeneous reactions with OH and HO₂ in their model and found
116 that the addition of these reactions increased the total ozone column abundance; the re-
117 sults were in better agreement with SPICAM total ozone column measurements than
118 the gas-only simulation, although ozone in their 3-D model was overpredicted at low lat-
119 itudes during the aphelion season (between $L_S = 60^\circ - 150^\circ$). However, Clancy et al.

120 (2016) found that the simulated ozone with the heterogeneous chemistry was not in agree-
 121 ment with MARCI total ozone column and, using the relationship between ozone and
 122 water ice as a proxy for heterogeneous reactions, found that the observed and simulated
 123 data had different correlations between ozone and water ice. They concluded that the
 124 observations did not show sufficient evidence of heterogeneous processes.

125 Lefèvre et al. (2021) used an adaptive semi-implicit scheme (ASIS) taken from Cariolle
 126 et al. (2017) to improve the chemical timestep and stability of the GCM. In contrast
 127 to previous studies, which had overestimated the water ice optical depth, this study had
 128 improved the water ice optical depth to be in better agreement with observations. These
 129 changes improved the agreement between simulated total ozone column and observed
 130 total ozone column from SPICAM at high northern latitudes. They found that the in-
 131 clusion of heterogeneous chemistry enhanced ozone abundance at high northern latitudes
 132 and was in strong agreement with SPICAM observations. The simulated water vapour
 133 was also in good agreement with total water vapour columns from SPICAM observations,
 134 although water vapour was overpredicted in some regions (e.g. low latitudes in aphelion,
 135 southern latitudes during southern summer). Water vapour is a key species to investi-
 136 gating the effect of heterogeneous chemistry since it is directly related to the abundance
 137 of HO_x . For Lefèvre et al. (2021), ozone abundance was still underpredicted in some sce-
 138 narios when water vapour abundance was low, but matched when water vapour abun-
 139 dance was higher ($> 1 \text{ pr-}\mu\text{m}$).

140 The adsorption of HO_x onto water ice is a physical process rather than chemical,
 141 and thus the method for modelling heterogeneous reactions is ambiguous as there are
 142 no definitive chemical reactions. Lefèvre et al. (2008) did not define any products for the
 143 heterogeneous reactions, and Lefèvre et al. (2021) later included oxygen and water vapour
 144 as products to conserve mass. Modelling of ozone has shown mixed results in explain-
 145 ing observed data and there are still many outstanding problems (Lefèvre et al., 2008;
 146 Clancy et al., 2016; Lefèvre et al., 2021). All studies regarding heterogeneous chemistry
 147 have used ozone total column abundance in their investigations, which has a major draw-
 148 back of only revealing the net difference in ozone throughout the column. In contrast,
 149 vertical profiles show the full vertical distribution of species (e.g. ozone and water ice)
 150 and, since the species can be viewed as mixing ratios rather than abundances, the re-
 151 lationship between ozone and water ice can be studied with equal weighting across all
 152 altitudes. Confirming the presence of ozone and water ice at the same altitudes is cru-
 153 cial to investigating the relationship between the two species as, without defining the ver-
 154 tical distribution, it is difficult to verify any impact heterogeneous reactions have on ozone.
 155 Determining the chemical impacts of heterogeneous reactions on both the total abun-
 156 dance and vertical distribution of ozone is essential to understanding how ozone is ex-
 157 pected to vary indirectly under the influence of heterogeneous reactions.

158 1.2 Outline

159 This study uses a combination of statistical analysis and 1-dimensional (1-D) mod-
 160 elling to quantify the impacts of heterogeneous reactions on ozone vertical profiles un-
 161 der different circumstances. We analyse observed vertical profiles of ozone and water ice
 162 at high ($> 45^\circ$) northern and southern latitudes between $L_S = 0^\circ - 180^\circ$ using a ver-
 163 tical cross-correlation analysis to determine the relationship between observed ozone and
 164 water ice. We then use a 1-D model with an improved heterogeneous chemical scheme
 165 to compare vertical profiles of simulated ozone with and without the heterogeneous chem-
 166 istry to determine the impact of heterogeneous reactions. We compare a high and a low
 167 water vapour abundance scenario to replicate the atmospheric state in the northern sum-
 168 mer and southern winter respectively, and assess the influence of heterogeneous chem-
 169 istry on ozone under such conditions.

170 Section 2 describes the vertical profiles from NOMAD, the vertical cross-correlation
 171 analysis, and the improved chemical scheme used in the 1-D model. Section 3 describes
 172 the results of the cross-correlation analysis between ozone and water ice, and the ozone
 173 variation in the 1-D model in low and high water vapour abundance scenarios. Finally,
 174 Section 4 discusses the impact of water ice on ozone, and the influence of water vapour
 175 on heterogeneous chemistry, before summarising the conclusions and implications of the
 176 study.

177 2 Methods

178 2.1 NOMAD Profiles

179 Ozone and water ice vertical profiles used in this study are derived respectively from
 180 the UVIS (UV-Visible) spectrometer and the SO (Solar Occultation) spectrometer on
 181 the NOMAD instrument aboard TGO, detailed in M. R. Patel et al. (2021) and Liuzzi
 182 et al. (2020). Data cover high latitudes ($> 45^\circ$ N/S) from $L_S = 0^\circ - 180^\circ$ MY 35 and,
 183 as there is little zonal variation, all longitudes are included together. Data between $L_S =$
 184 $0^\circ - 30^\circ$ are omitted for the high northern latitudes analysis due to an absence of data
 185 below 30 km in ozone and water vapour profiles. Between $L_S 30^\circ - 180^\circ$ at the high north-
 186 ern latitudes, there is a much lower ozone abundance (up to 0.1 ppmv) and thus there
 187 are fewer ozone profiles which meet the minimum requirement for the analysis (Table
 188 1). This still leaves 300 profile pairs which meet the conditions set for the cross-correlation.

189 Solar occultations profile the atmosphere at the terminator and can occur up to
 190 24 times a sol (martian day) due to the nature of the orbit of the spacecraft. Ozone is
 191 retrieved within the Hartley Band between wavelengths 240 – 320 nm, while water ice
 192 is retrieved using five diffraction orders which range between 2.2 μm and 4.3 μm . See M. R. Pa-
 193 tel et al. (2021) and Liuzzi et al. (2020) for the full retrieval process for ozone and wa-
 194 ter ice respectively.

195 2.2 Vertical Cross-Correlation

196 We conduct a cross-correlation between ozone and water ice vertical profiles retrieved
 197 from the NOMAD instrument to assess the vertical relationship between the two species.
 198 In terrestrial studies, this technique is often used between two time series to determine
 199 whether there is a correlated time lag between the variables, along with the nature of
 200 this time displacement (e.g. Arattano and Marchi (2005); Peppas et al. (2017)). It is of-
 201 ten used when one variable is expected to influence the other and there may be a delay
 202 in the response. By displacing the altitude, any vertically displaced patterns between
 203 the ozone and water ice can be identified, which, due to the suppression of HO_x in water-
 204 ice clouds, may impact ozone at different altitudes. In addition, water-ice clouds can span
 205 several kilometres and, with a standard Pearson’s correlation, any increase in ozone which
 206 does not span the full altitude range of the water-ice cloud would not be consistently de-
 207 tected. By conducting a cross-correlation, any variation in ozone within the water-ice
 208 cloud may be detected within a few kilometres displacement. Furthermore, a standard
 209 correlation may not detect any relationship if the vertical profiles of water ice and ozone
 210 are displaced at differing altitudes due to any interpolation error, while the cross-correlation
 211 is able to capture this with the vertical lag.

212 Each cross-correlation conducted on a pair of vertical profiles produces multiple cor-
 213 relations at different altitude displacements. In this study, a 2-tailed Student’s T-test
 214 is used to assess the p-value of each correlation at a significance level of 5% and, from
 215 all the significant correlations, the one with the lowest p-value is selected as the most
 216 significant correlation. The p-value is the probability of a result being at least as extreme
 217 as the observed datapoint, with the assumption that the null hypothesis is true. In this
 218 analysis, it is the probability of obtaining a correlation at least as extreme as the observed

Table 1. Total number of vertical profiles for high northern and southern latitudes, followed by the number of profiles which both have at least 1 ppmv of water ice, 0.03 ppmv of ozone, and a minimum of 6 datapoints in each profile. Positive correlations are only included if the altitude displacement is within ± 10 km.

Latitude ($^{\circ}$)	Total Profiles	Profiles used	Positive correlations	Percentage positive
≥ 45 N	383	249	90	36.1%
≥ 45 S	711	564	314	55.7%

219 correlation between ozone and water ice, given the assumption that there is no vertical
220 correlation between ozone and water ice.

221 Profiles with no variation and very low volume mixing ratios (vmr) are unsuitable
222 for the vertical correlation analysis as they can produce false correlations that would ob-
223 scure results in the rest of the analysis. Including such profiles may return positive cor-
224 relations between ozone and water ice, which have no physical meaning as the vmr is so
225 low. As a result, a minimum condition of at least one value in each ozone profile exceed-
226 ing a threshold of 0.03 ppmv is set. This restriction enables the vertical variation of ozone
227 in regions of low abundance (such as in the northern hemisphere) to be included in the
228 analysis. Panel (b) of Figure 1 shows the ozone vertical profiles $> 45^{\circ}$ N for $L_S 0^{\circ} -$
229 180° MY 35. Panels (c) and (d) show the vertical profiles of water ice and water vapour
230 respectively at the same times and locations, while panel (a) gives the corresponding lat-
231 itude and local time of the profile. Panels (e), (f), (g), and h show the same species but
232 for $> 45^{\circ}$ S. Similarly, a minimum threshold of at least one value in each water ice pro-
233 file exceeding 1 ppmv is used. These values are defined graphically and the sensitivity
234 analysis undertaken indicated that the number of profiles retained for the analysis were
235 robust for the threshold (see Supplementary Information for the sensitivity analysis). In
236 addition, a threshold of a minimum of 6 datapoints in each profile is used to filter pro-
237 files with poor vertical coverage and which otherwise would be unsuitable for the cross-
238 correlation analysis and lead to insignificant correlations.

239 Figure 2 shows an example of the cross-correlation on a pair of profiles; the max-
240 imum correlation between the profiles occurs when the ozone is 4 km above the water
241 ice (red cross on right panel; correlation 0.85). This corresponds to the increase in ozone
242 between 34–28 km, peaking at 0.45 ppmv matching with the increase in water ice be-
243 tween 31–25 km which peaks at 7.0 ppmv (left panel). At an altitude displacement $<$
244 -10 km the decrease in ozone from the peak between 28–24 km matches with the de-
245 crease in water ice which occurs from 25 – 10 km, causing the correlation at this dis-
246 placement to increase to 0.8. See the Supplementary Information for further details on
247 the methodology used to calculate the cross-correlation and filter data.

248 2.3 1-D Model

249 The 1-D model used in this work (henceforth referred to as the 1-dimensional Mar-
250 tian Photochemical Model, 1-D MPM) is derived from the Open University modelling
251 Group Mars GCM (MGCM), which exists as a collaboration between the Laboratoire
252 de Météorologie Dynamique (LMD), the Open University, the University of Oxford, and
253 Instituto de Astrofísica de Andalucía (Forget et al., 1999).

254 By using a 1-D model, the interactions between tracers can be identified purely as
255 chemical interactions without the added complexity of a 3-D dynamical model, such as
256 the transport of heat, chemical species, and aerosols. The effects of heterogeneous re-

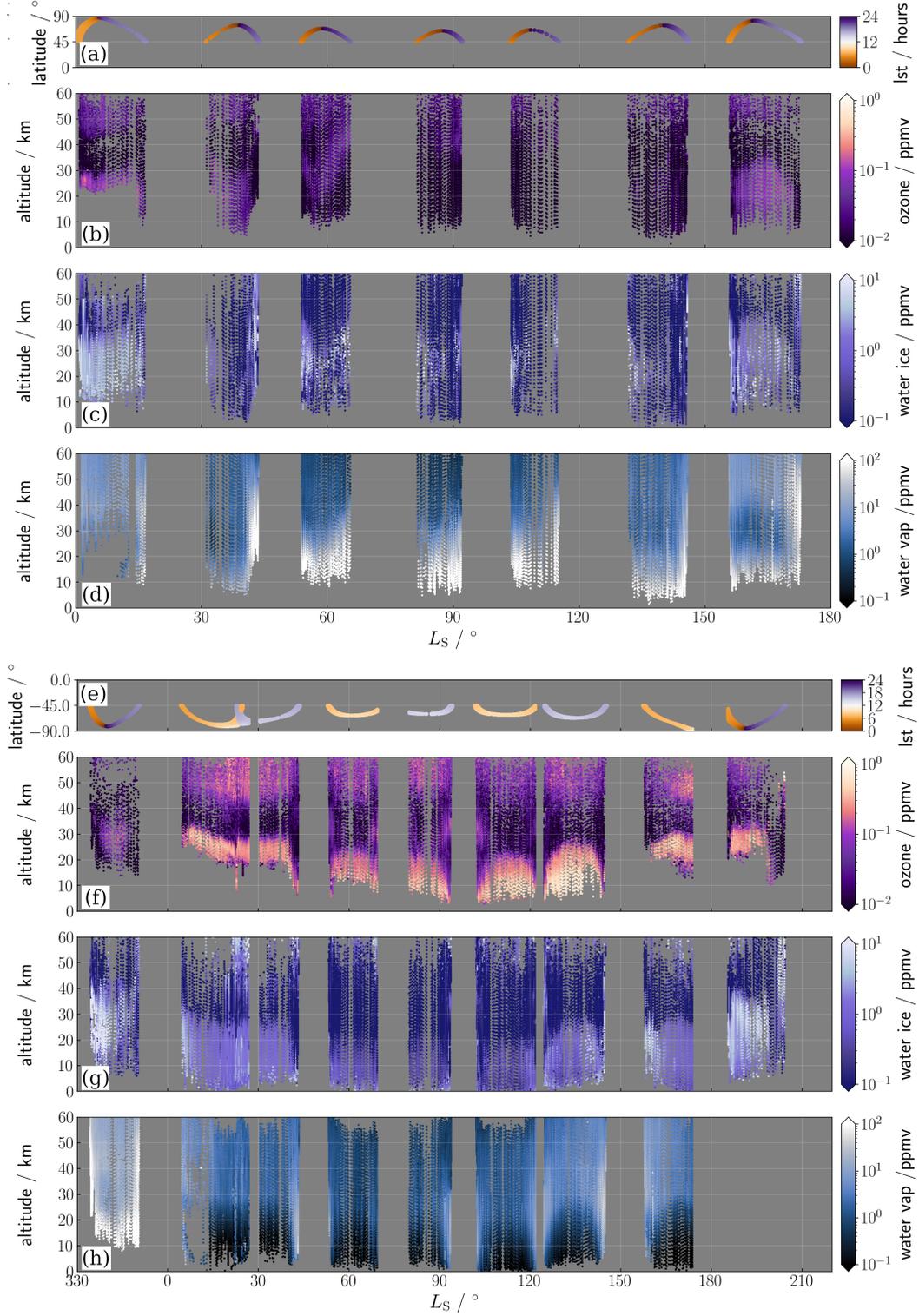


Figure 1. (a) The latitude and local time at which each observed profile is taken. Vertical profiles at high northern latitudes ($> 45^\circ$ N) of (b) ozone (M. R. Patel et al., 2021); (c) water ice (Liuzzi et al., 2020); and (d) water vapour (Villanueva et al., 2022) from $L_S = 0^\circ - 180^\circ$ MY 35 from the NOMAD instrument. (e), (f), (g) and (h); equivalent as above but for high southern latitudes ($> 45^\circ$ S).

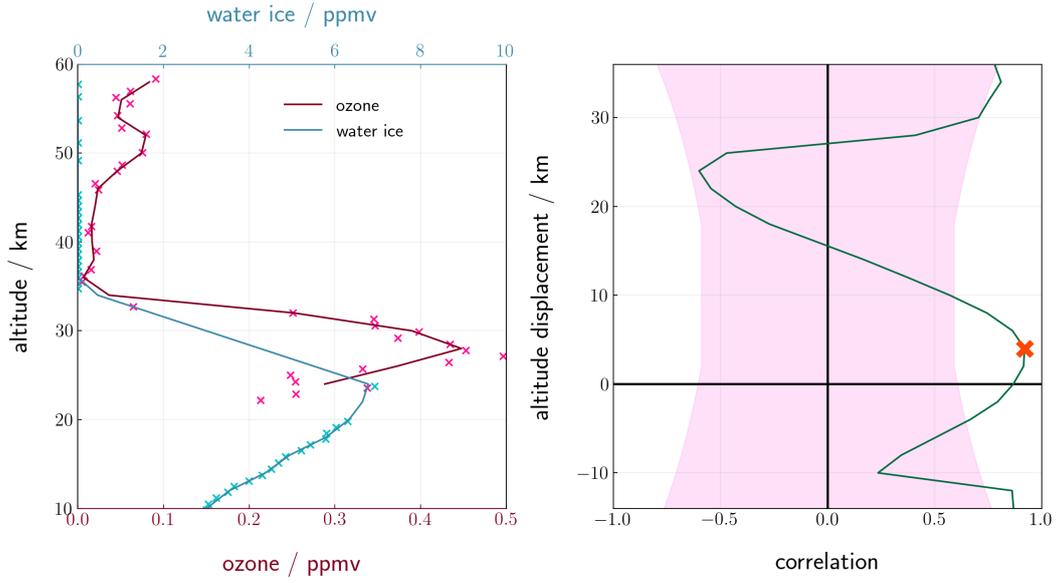


Figure 2. An example of a cross-correlation analysis between a pair of profiles: (left panel) vertical profiles of (blue) water ice; and (red) ozone, where the crosses are the datapoints and the lines are the interpolated profiles every 2 km: (right panel) the (green) vertical correlation between the profiles across all displacements. A positive displacement of 4 km corresponds to a shifting of the water ice profile up by 4 km. The pink area marks the correlations which are not significant according to a Student's T-test at 5% significance. The red cross indicates the most significant vertical correlation of the cross-correlation. Profiles are from NOMAD/TGO at $L_S = 187^\circ$, latitude 70.5° S, 2020.04.21 11:04:57 UTC.

257 actions on ozone can therefore be isolated and, by running the model both with and with-
 258 out the heterogeneous reactions, the difference between ozone in both scenarios highlights
 259 the direct impact of heterogeneous chemistry on ozone.

260 The 1-D MPM is compiled with 70 levels, spaced non-linearly according to pres-
 261 sure, with 22 tracers and 60 chemical and photochemical reactions. Initial starting
 262 conditions are taken from the outputs of the MGCM, run with data assimilation of temper-
 263 ature and dust retrievals (J. A. Holmes et al., 2020) from the Mars Climate Sounder (MCS)
 264 instrument aboard the Mars Reconnaissance Orbiter, with photochemistry and the new
 265 ASIS chemical scheme. The heterogeneous chemical scheme, originally taken from Lefèvre
 266 et al. (2021), has been updated to improve the representation of heterogeneous chem-
 267 istry. The products of the improved heterogeneous reactions are treated as a separate
 268 species rather than recycled into water vapour and oxygen. The improved scheme also
 269 includes a third heterogeneous reaction using H_2O_2 (Pouvesle et al., 2010). The OH, HO_2
 270 and H_2O_2 are converted into three separate species which are unable to react in any way
 271 aside from converting back to their original species. In this heterogeneous scheme, wa-
 272 ter ice is treated as a sink for HO_x . For OH, this is represented by,



273 where ice_{OH} is the added species, representing a ‘sink’ for the OH which is adsorbed
 274 onto water-ice particles. The concentration of the ice_{HO_x} (and $\text{ice}_{\text{H}_2\text{O}_2}$) species only de-
 275 creases when either the tendency of water ice (the rate of change from the previous timestep)
 276 is negative and thus water ice sublimates to water vapour or if the water ice abundance
 277 is 0.

278 Unlike for OH and HO_2 , the adsorption of H_2O_2 onto water ice is a reversible re-
 279 action (Pouvesle et al., 2010). The Langmuir adsorption model explains the reversible
 280 reaction of the adsorption of species on a monolayer surface. Therefore, the Langmuir
 281 reaction rate, which has been derived from laboratory experiments, is used for the het-
 282 erogeneous reaction rate of H_2O_2 in the model:

$$\theta_A = \frac{K_{eq}^A p_A}{1 + K_{eq}^A p_A}, \quad (2)$$

283 where p_A is the partial pressure of the species (e.g. HO_x), K_{eq}^A is the equilibrium
 284 constant of the reaction, and θ_A is the fractional coverage, or the ratio between the vol-
 285 ume of gas adsorbed onto the surface to the volume of gas adsorbed onto the surface at
 286 maximum occupancy, assuming only monolayer of adsorbate onto the surface (Langmuir,
 287 1918).

288 Figure 3 shows the vertical profile of ozone for the previous and new heterogeneous
 289 schemes and the differences between them over one sol (note the non-linearity on the dif-
 290 ference colourbar on the bottom panel). At lower altitudes < 25 km there is little dif-
 291 ference in ozone abundance between the simulations (indicated by the pale green in the
 292 bottom panel). However, at higher altitudes, the new heterogeneous scheme simulates
 293 a lower ozone abundance throughout the sol, between 30–55 km (red in the bottom panel).
 294 This is likely due to the water ice, which formed at a uniform concentration between 30–
 295 50 km throughout the sol in both simulations (not shown). During the day (between 0600–
 296 1800 hours), ozone abundance in the new scheme gradually increases up to 0.1 ppmv, and
 297 the altitude range it spans also increases from 30–40 km to 30–60 km (top panel). In con-
 298 trast, ozone has little variation during the day in the old heterogeneous scheme, and has
 299 a higher abundance (0.1 ppmv between 0600–1800 hours), which ranges between 30–45 km.
 300 The greatest difference between the two schemes occurs at nighttime, where the ozone
 301 abundance in the new scheme increases to 0.8 ppmv around 50 km, while in the old scheme,

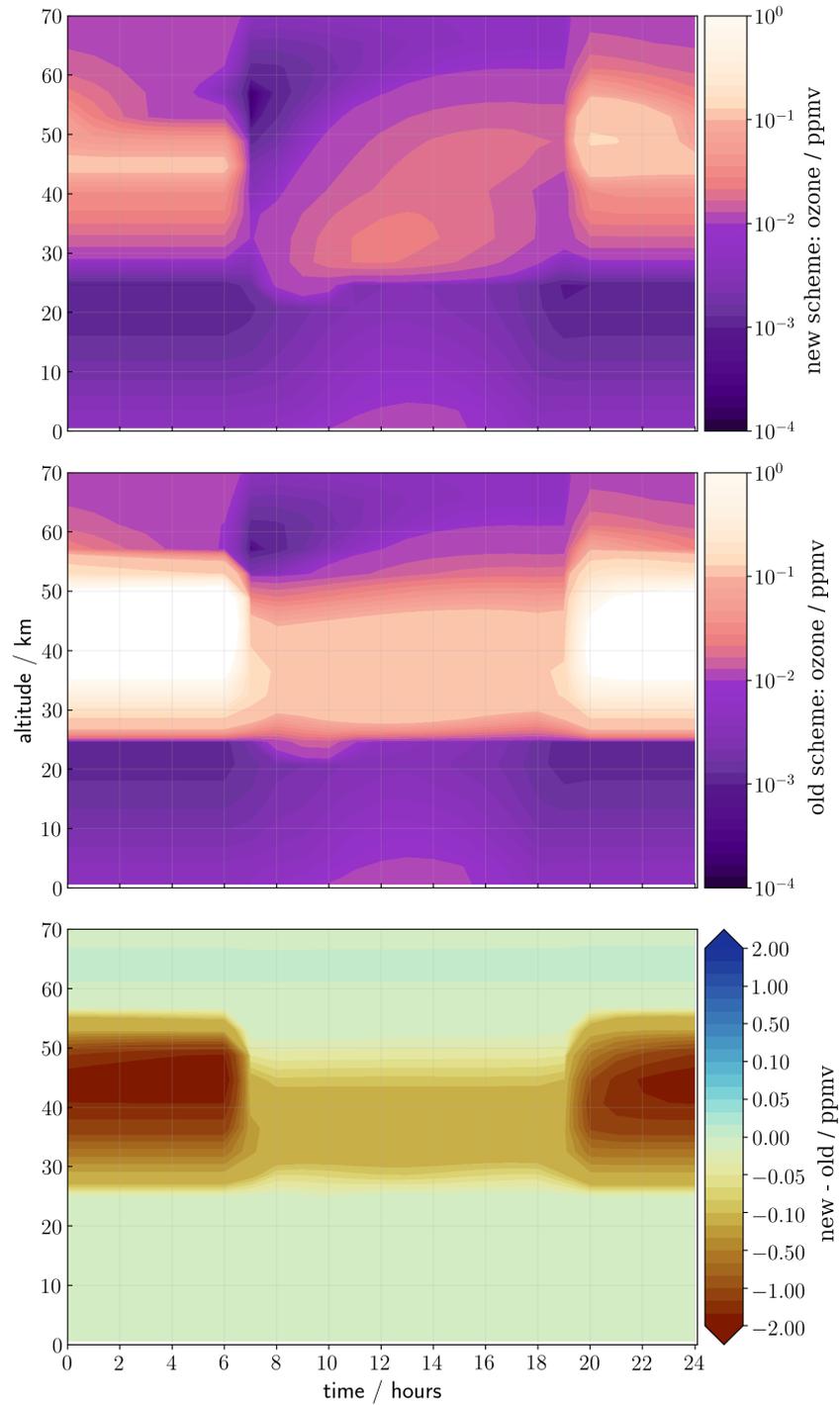


Figure 3. Simulation of ozone abundance over a sol at 0° S, $L_S = 15^\circ$ with the 1-D MPM using; (top) the new heterogeneous chemistry scheme; (middle) the old heterogeneous chemistry, and (bottom) the difference between the two.

ozone increases to > 2 ppmv between 30–50 km. Overall, the new heterogeneous scheme has a lower ozone total column abundance, due to the differences > 30 km.

Two scenarios with a high and low water vapour abundance are simulated with the 1-D MPM to represent high northern and southern latitudes respectively; the scenarios are a simplified representation of high latitude regions and longer temporal periods during the aphelion season. To reduce the parameter differences between the high and low water vapour scenarios, the latitudes and local times are kept consistent at latitude 0° and 1200 local solar time (LST); only the time of year is changed, with the low and high water vapour scenario at $L_S = 60^\circ$ and $L_S = 180^\circ$ respectively. By varying the L_S , the solar insolation decreases in the $L_S = 60^\circ$ simulation and results in a lower water vapour abundance due to lower temperatures. The vertical structure of temperature remains similar in both scenarios due to the same latitude used, and the local solar time ensures it is a photochemically active part of the day. This results in the water-ice clouds form at similar altitudes in both scenarios.

Despite the vertical profiles covering high latitudes, the model scenarios are used only to investigate the ozone variation under different water vapour abundance and not the latitudes themselves. Due to the limitations in the condensation scheme of the model, it is not appropriate to run the 1D-MPM at high latitudes, and thus only low latitudes are used in this study. For verification, the 1-D MPM was run at different L_S , local time, and latitudes to investigate the chemical response of ozone from the heterogeneous reactions. The two scenarios used in this study produce similar results to the other simulations and are representative of the ozone variation (not shown). The results from these runs were verified against the MGCM from J. Holmes et al. (2021) and the JPL Caltech 1-D model from Viúdez-Moreiras et al. (2019) and found to be in good agreement (see Supplementary Information for a full comparison).

The model is run both with and without heterogeneous chemistry, with the gas-phase only simulation used as a control to investigate the chemical impacts from the heterogeneous reactions. Further details of the adapted heterogeneous chemistry parameters of the model are in the Supplementary Information.

3 Results

The two high latitude regions ($> 45^\circ$ N/S) used in this study cover $L_S = 0^\circ - 180^\circ$ MY 35 and have contrasting abundances of water vapour. The time period incorporates the northern spring equinox to the northern autumnal equinox, where, from previous studies, ozone is underpredicted. Figure 4 shows the averaged water ice profiles $> 45^\circ$ S between $L_S = 100^\circ - 120^\circ$ and the ozone difference between modelled and observed profiles. An ozone difference greater than 0 ppmv (dark red bars) indicates an underprediction in the MGCM, which follows a similar trend to the water ice abundance. The greatest ozone deficit occurs at lower altitudes (< 20 km) and coincides with a higher water ice concentration, while a small deficit occurs above 40 km. Figure 4 is representative of the relationship between the ozone deficit and water ice present throughout the southern hemisphere from $L_S = 0^\circ - 180^\circ$ (not shown).

During northern summer, the water ice cap melts, leading to a high abundance of water vapour in the northern hemisphere (Clancy & Nair, 1996; Montmessin & Lefèvre, 2013; Steele et al., 2014). Using the observed vertical profiles from NOMAD between 10–50 km, the total water vapour abundance is, on average, one magnitude larger in the northern hemisphere than the southern (1.37×10^{-2} pr- μm compared to 1.78×10^{-3} pr- μm) [From Geronimo 2022]. Note that not all occultations extend to the surface, and thus the calculated total column abundance only extends down to 10 km. This is demonstrated in panel (d), Figure 1, where between $L_S = 30^\circ - 180^\circ$ the water vapour reaches concentrations > 100 ppmv at 30 km and below. The total column of water vapour is there-

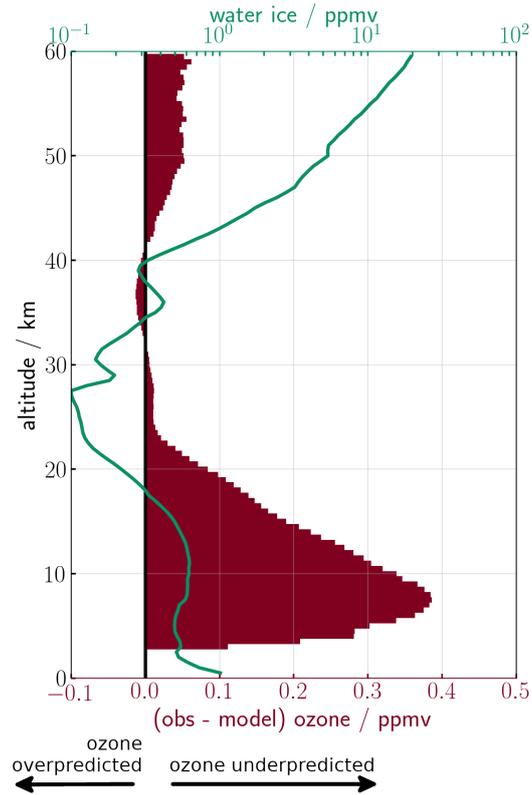


Figure 4. Averaged profiles (212 vertical profiles) $> 45^\circ$ S between $L_S = 100^\circ - 120^\circ$ of (green line) observed water ice, and (red bars) the difference in ozone (observed – modelled). Positive values indicate an underprediction of ozone. Modelled data are taken from the MGCM, assimilated with temperature and dust retrievals from MCS.

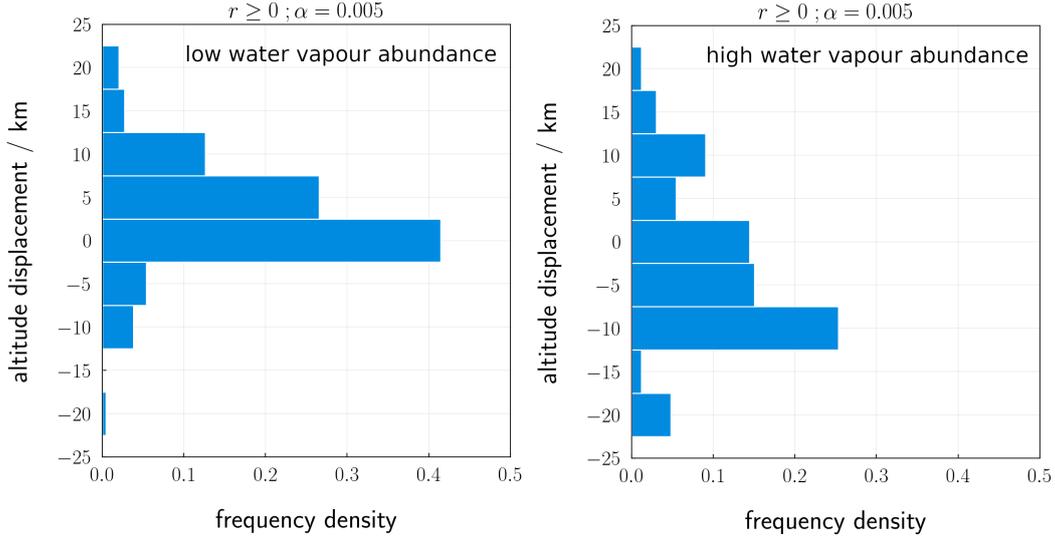


Figure 5. Histogram of altitude displacements for significant, positive, vertical correlations between ozone and water ice observed profiles for (left) southern latitudes ($> 45^\circ$ S) $L_S = 0^\circ - 180^\circ$ and (right) northern latitudes ($> 45^\circ$ N) between $L_S = 30^\circ - 180^\circ$. α shows the significance level, and $r \geq 0$ indicates positive correlations only. Bins cover 5 km, with the y axis label denoting the centre of the bin (e.g. the 0 km bin includes values from 2.5 km to 2.5 km.)

352 fore likely much higher, as the lowest altitudes have the greatest contribution to the to-
 353 tal column measurement.

354 3.1 Low Vapour Case: Ozone – Water Ice Relationship

355 The results from cross-correlation between the ozone and water ice vertical pro-
 356 files in the southern latitude region suggest there are significant, positive vertical cor-
 357 relations when the ozone and water ice profiles match at the same altitudes. The absence
 358 of high abundances of water vapour in the southern hemisphere allows the relationship
 359 between ozone and water ice to be seen without interference from the presence of wa-
 360 ter vapour. The cross-correlation results demonstrate a clear relationship between ozone
 361 and water ice within a vertical displacement range of ± 10 km of the two species; the left
 362 panel of Figure 5 shows a histogram of the altitude displacement from the most signif-
 363 icant positive correlations of each profile pair in the southern latitude region. An alti-
 364 tude displacement of 4 km indicates a positive shift in the water ice profile from its orig-
 365 inal position by 4 km, and thus indicates that the water ice profile matches with an ozone
 366 profile which is 4 km above. An altitude displacements of 0 km implies a correlation be-
 367 tween the two species at their original position.

368 In the histogram, there is a sharp peak between -2.5–2.5 km with a frequency den-
 369 sity of 0.41, implying that water ice and ozone profiles correlate strongly at the same alti-
 370 tude, or when there is a small altitude displacement (< 2.5 km). Table 1 shows the num-
 371 ber of profiles used in the analysis and the percentage of profiles which have positive cor-
 372 relations within ± 10 km. Of the profiles suitable for cross-correlation, 55.7% of the most
 373 significant correlations are positive within this altitude range.

374 The positive correlation between ozone and water ice is continuous throughout the
 375 season can be seen visually in panels (f) and (g) in Figure 1. Both species follow a sim-

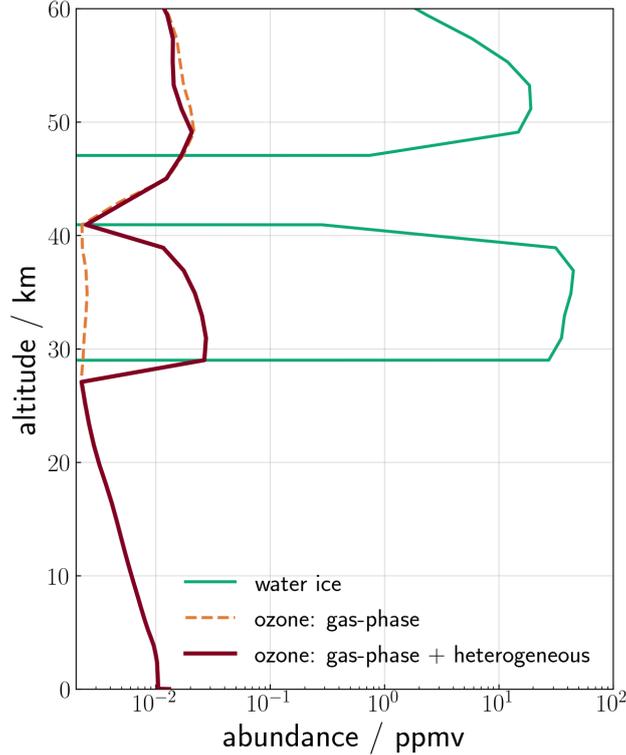


Figure 6. A single modelled vertical profile from the 1-D MPM of (dark red) ozone from the heterogeneous run, (dashed orange) ozone from gas-phase only run and (green) water ice. Profiles are simulated at 1200 LST, latitude 0° , $L_S = 0^\circ$.

376 ilar latitudinal trend with abundances ranging between 0.2–1 ppmv for ozone and 0.5–
 377 3 ppmv for water ice between 0–30 km. The cross-correlation matches the variation in
 378 ozone and water ice profiles together at 0 km displacement, as both species have a sim-
 379 ilar vertical variation. In panels (f) and (g) of Figure 1, both the water ice and ozone
 380 have increased abundances below 30 km, which decrease above this altitude.

381 During this period, the water vapour abundance is low due to the cold atmospheric
 382 temperatures during southern winter. Panel (h) of Figure 1 shows the vertical profiles
 383 of water vapour derived from NOMAD observations from (Villanueva et al., 2022); the
 384 water vapour ranges from < 0.1 ppmv to 30 ppmv, with much of the water vapour only
 385 observed at altitudes above 30 km.

386 Figure 6 shows a modelled profile from the 1-D MPM both with and without het-
 387 erogeneous chemistry taken at 1200 LST, $L_S = 0^\circ$, latitude 0° as explained in Section 2.
 388 Between 30–40 km there is an increase in ozone in the heterogeneous simulation (dark
 389 red) which is not observed in the gas-phase-only simulated ozone (dashed orange). This
 390 increase is due to the heterogeneous reactions which occur when water ice (green) is present.
 391 The water ice adsorbs the HO_x and reduces the HO_x abundance, which results in a lower
 392 ozone destruction rate and enhances the ozone abundance. The peak increase in ozone
 393 from heterogeneous reactions in the low water vapour abundance simulations are pro-
 394 portional to 43 – 75% of the peak ozone deficit shown in Figure 4.

395 The ozone simulated with heterogeneous chemistry between 50–60 km does not
 396 exhibit the same response, however, despite the formation of water ice at these altitudes.
 397 Indeed, the ozone simulated with heterogeneous chemistry is slightly lower than the gas-

398 phase-only ozone. This is due to a decrease in water ice abundance between 50–60 km
 399 from the previous timestep, which causes the release of previously adsorbed hydroxyl rad-
 400 icals, increasing their abundance compared to the control (gas-phase-only) simulation.

401 3.2 High Vapour Case: No Ozone – Water Ice Correlation

402 Given that we have found that the altitude displacement between ozone and wa-
 403 ter ice is likely with ± 10 km altitude range in the southern hemisphere, exploring the cor-
 404 relation between ozone and water ice in the northern hemisphere at this altitude range
 405 could yield similar results. The right panel of Figure 5 shows the histogram of signif-
 406 icant, positive vertical correlations for observed ozone and water ice profiles at high north-
 407 ern latitudes. Unlike in the southern hemisphere, the distribution of altitude displace-
 408 ments in the northern hemisphere is more uniform and there is no clear correlation be-
 409 tween ozone and water ice in one altitude displacement range. In addition to this, there
 410 are proportionally fewer significant positive correlations, with only 36.1% of the most
 411 significant correlations being positive within ± 10 km displacement, compared to 55.7%
 412 in the southern latitude region (see Table 1).

413 As discussed previously, the high northern latitudes between $L_S = 30^\circ - 180^\circ$ have
 414 a higher water vapour abundance than the high southern latitudes (lower panels of Fig-
 415 ure 1), as the time period covers the northern summer and the sublimation of the water-
 416 ice cap (Montmessin & Lefèvre, 2013; Steele et al., 2014). To simulate similar conditions
 417 of higher water vapour abundance, a 1-D MPM simulation at latitude 0° , $L_S = 180^\circ$
 418 is used to investigate the response of ozone to the heterogeneous reactions. The verti-
 419 cal profile from this model run is then compared to another with a much lower water vapour
 420 abundance at $L_S = 60^\circ$; total column water vapour abundance is $9.3 \text{ pr-}\mu\text{m}$ and 3.8 pr-
 421 μm for the high and low vapour scenarios respectively. Note that these values are the
 422 total column abundance for the full vertical column and thus not directly comparable
 423 to the abundances calculated for the observed vertical profiles which only include alti-
 424 tudes between 10–50 km. The total column water vapour abundances between 10–50 km
 425 are $2.1 \text{ pr-}\mu\text{m}$ and $0.008 \text{ pr-}\mu\text{m}$ for the high and low water vapour scenario respectively.

426 Figure 7 demonstrates the impact of heterogeneous chemistry on ozone abundance
 427 for low ($L_S = 60^\circ$; panels (a) and (b)), and high ($L_S = 180^\circ$; panels (c) and (d)) wa-
 428 ter vapour abundances through the change in hydroxyl radicals. Panels (a) and (c) show
 429 the vertical profile of water ice and water vapour (lines) as well as the ozone residual (bars)
 430 which is the heterogeneous simulation minus the gas-phase simulation. Panels (b) and
 431 (d) shows the vertical profile of ozone and HO_x (combined OH + HO_2). Panel (a) shows
 432 an excess of ozone (dark red bars) from the heterogeneous simulation up to 0.05 ppmv
 433 in the low water vapour scenario (shown by the water vapour profile; dark blue) between
 434 25–40 km. Panel (b) shows the ozone (dark red) and HO_x (OH and HO_2 ; orange) abun-
 435 dance for the control (solid) and heterogeneous (dashed) simulation; the large difference
 436 between the two HO_x abundances (over two orders of magnitude difference between 25–
 437 30 km) is a direct result of the heterogeneous reactions caused by the presence of water
 438 ice at these altitudes. In contrast, panel (c) only has an ozone residual of up to 0.005 ppmv
 439 in the high water vapour abundance scenario despite having a larger water ice abundance
 440 (light blue) of 30 ppmv compared to 10 ppmv. The HO_x abundance is over an order of
 441 magnitude higher in the control simulation between 30–45 km and the adsorption of HO_x
 442 onto water ice is proportionally much smaller than the lower water vapour scenario; be-
 443 tween 30–35 km, HO_x decreases from 3×10^{-3} to 1×10^{-3} ppmv, while in the low vapour
 444 scenario HO_x decreases from 3×10^{-4} to 1×10^{-5} ppmv despite the water ice abun-
 445 dance being three times less.

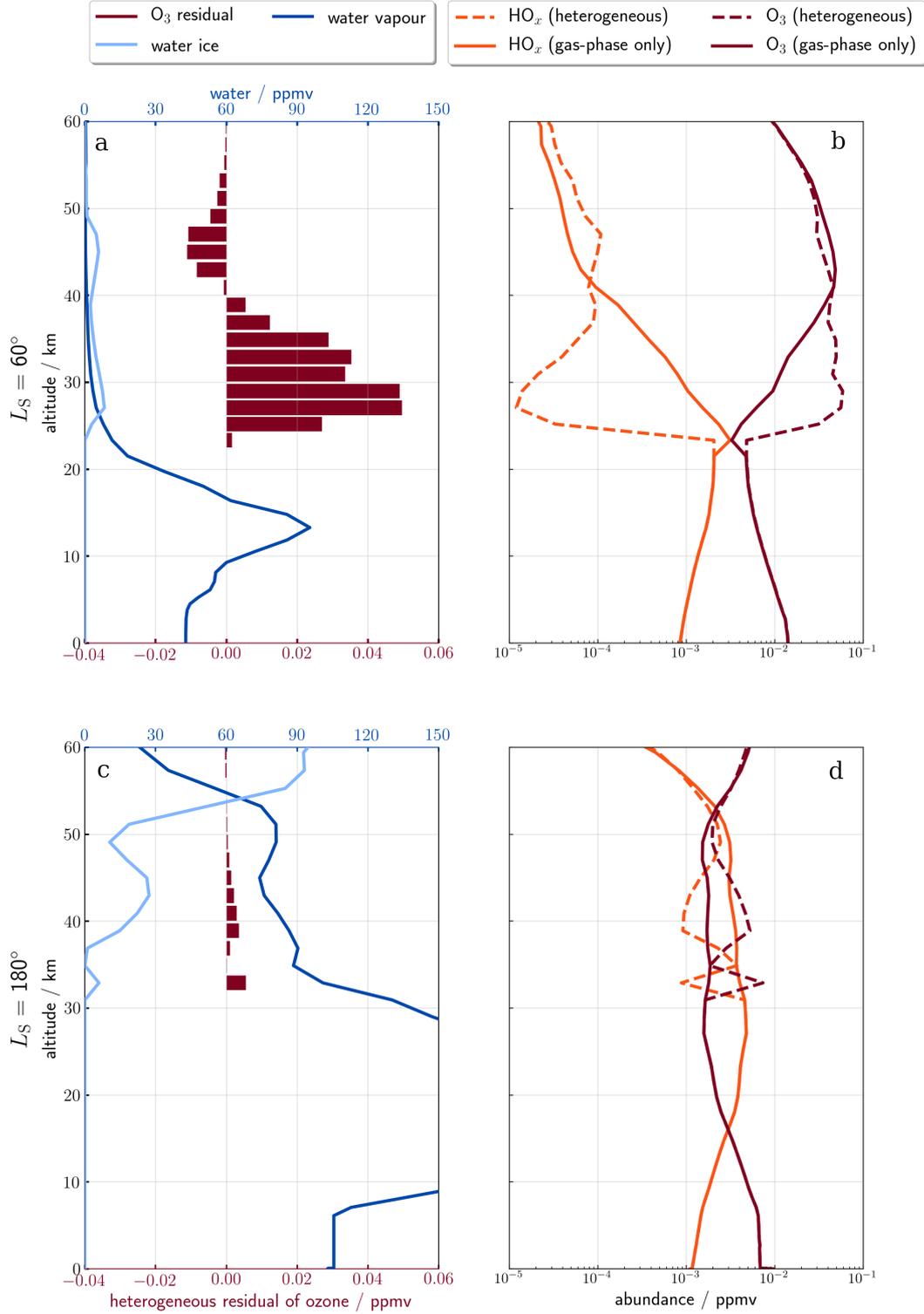


Figure 7. Modelled profiles from the 1-D MPM of (a and b) low water vapour (at $L_S = 60^\circ$), and (c and d) high water vapour (at $L_S = 180^\circ$) at latitude 0° , local solar time 1200 hours. (First column; (a) and (c)) vertical profiles of (light blue) water ice, (dark blue) water vapour, and (dark red bars) the ozone residual (calculated by subtracting the heterogeneous ozone from the gas-phase only ozone). Abundance difference for ozone is on the bottom x-axis and abundance for water ice and vapour is on the top x-axis. (Second column; (b) and (d)) vertical profiles of (dark red) ozone and (orange) HO_x for (dashed) heterogeneous and (solid) gas-phase-only simulations. Note the abundances are on a logarithmic scale.

4 Discussion

4.1 Impact of Heterogeneous Chemistry on Ozone

The positive vertical correlation between ozone and water ice in the southern latitude region can be used as a proxy for the presence of heterogeneous reactions. The results of the correlation analysis are supported by the 1-D MPM simulations which, in Figure 6, demonstrate the difference between ozone profiles in the presence of water ice, both with and without heterogeneous chemistry. The increase in ozone vmr due to heterogeneous reactions occurs at the same altitudes water ice is present in. According to the simulated profile in Figure 6, this is between 30 – 40 km.

The increase in ozone vmr due to heterogeneous reactions occurs at the same altitudes as water ice is present, which, in the simulated profile in Figure 6, is between 30–40 km. The heterogeneous reactions act as a sink for HO_x , reducing the abundance of HO_x as the water-ice surface area increases. As HO_x reactions with ozone are one of the main destructive pathways of ozone, the decrease in HO_x vmr ultimately leads to an increase in ozone vmr within the altitudes that water ice has formed.

The adsorption rate of heterogeneous reactions is dependent on the water-ice surface area, HO_x molecular density, and temperature. At higher altitudes, all three factors are typically lower due to the exponential decrease in pressure and general decrease in temperature. The impact of heterogeneous chemistry, therefore, is reduced at higher altitudes. The decrease in molecular density of HO_x with increasing altitude results in a smaller difference between the simulations, with and without heterogeneous chemistry; this is shown by the decrease in difference of ozone, as altitude increases in Figures 6 and 7. This suggests that heterogeneous chemistry has less influence on the ozone abundance at high (> 50 km) altitudes, which is reflected in the ozone deficit between simulated and observed vertical profiles as seen in Figure 4. The largest ozone deficit between simulated and observed profiles coincides with a higher water ice concentration below 30 km, while only a small deficit occurs above 40 km. The positive vertical correlation between ozone and water ice occurs below 30 km; this relationship matches the response observed in the 1-D MPM heterogeneous simulation, in which ozone increases at the same altitudes water ice forms. Heterogeneous chemistry could therefore explain some (43–75%) of the ozone deficit, as the addition of the heterogeneous reactions increases ozone at the altitudes water ice is present and has a greater impact on ozone abundance at low altitudes. As the heterogeneous reactions do not fully account for the ozone deficit, it is possible that there are other unknown reactions occurring, or the rates of the heterogeneous reactions themselves may be incomplete.

4.2 Water Vapour Influence on Heterogeneous Chemistry

The lack of any positive vertical correlation between ozone and water ice at a consistent altitude range in the northern hemisphere suggests there is no relationship between ozone and water ice and, by proxy, no heterogeneous chemistry present. This conflicts with the clear relationship observed between the two species in the southern hemisphere.

One suggestion for this could be that, globally, there is no heterogeneous chemistry occurring and the relationship observed in the southern hemisphere is simply an anti-correlation between ozone and water vapour portrayed through the water ice distribution which, itself, appears to be an inverse of the water vapour distribution (Figure 1). If this were the case, however, the relationship observed in the northern hemisphere should be similar to that of the southern hemisphere.

These contrasting results between the high northern and southern latitudes therefore require an additional explanation. The 1-D MPM simulations of a high and low wa-

495 ter vapour scenario enable an explicit investigation into the chemical impact of the het-
 496 erogeneous reactions between HO_x and ozone. Despite the decrease in HO_x abundance
 497 from the control to the heterogeneous simulation in the high water vapour scenario (lower
 498 panels, Figure 7), there is little difference in ozone abundance, indicating that the het-
 499 erogeneous reactions have relatively little impact on ozone abundance in circumstances
 500 of high HO_x abundance. As HO_x is a by-product of water vapour photolysis, a high abun-
 501 dance of water vapour can be assumed proportional to a high HO_x abundance. In the
 502 northern hemisphere during northern summer, the water vapour abundance is higher than
 503 observed in the southern hemisphere and, by extension, the HO_x abundance is also likely
 504 to be greater. The relative decrease of HO_x abundance as a result of heterogeneous chem-
 505 istry has a negligible impact on the abundance of ozone and is unlikely to be detectable
 506 through ozone variation. Therefore, ozone and water ice would not be expected to have
 507 a positive relationship in scenarios of high water vapour abundance, as heterogeneous
 508 chemistry is offset by the additional availability of HO_x and the ozone abundance does
 509 not significantly increase under such circumstances.

510 4.3 Conclusions

511 We find that the influence of heterogeneous reactions on ozone is dependent on the
 512 abundance of water vapour, which undergoes seasonal and spatial variation. The rela-
 513 tionship between the observed ozone and water ice, which is used as a proxy for hetero-
 514 geneous chemistry, is also expected to vary temporally and spatially.

515 In the cross-correlation analysis at high southern latitudes, there is a positive verti-
 516 cal correlation between ozone and water ice at 0 km displacement. In contrast, in the
 517 northern latitudes, when there is a much higher water vapour abundance, there is no clear
 518 positive correlation. In the 1-D MPM simulations, the ozone abundance is much lower
 519 in the high water vapour scenario and the ozone residual between the heterogeneous and
 520 gas-phase simulation is closer to 0 ppmv, implying that there is a minimal effect (< 0.005 ppmv
 521 increase) of heterogeneous chemistry on ozone under high water vapour conditions.

522 The photolysis of water vapour is not the sole factor in the water cycle affecting
 523 the global ozone distribution — water ice also appears to influence ozone abundance in-
 524 directly through the adsorption of hydroxyl radicals.

525 The cross-correlation conducted on the ozone and water ice vmr profiles shows that
 526 these two variables are positively correlated in the vertical. This is contrary to previous
 527 studies which found an anti-correlation between the total abundance of ozone and wa-
 528 ter ice (Clancy et al., 2016). From investigating the ozone variation between 1-D MPM
 529 outputs with and without heterogeneous chemistry, we show that heterogeneous reac-
 530 tions between HO_x and water ice are a plausible explanation for the positive vertical cor-
 531 relation between ozone and water ice.

532 Furthermore, heterogeneous chemistry increases the ozone vmr at altitudes water
 533 ice forms, and thus could explain some of the ozone deficit in GCMs when compared to
 534 observations, which is in agreement with Lefèvre et al. (2008, 2021). While the addition
 535 of heterogeneous chemistry may not fully resolve the ozone deficit seen in GCMs, it in-
 536 creases the ozone abundance at locations which currently underpredict ozone, while hav-
 537 ing a minimal effect on ozone at higher altitudes and in areas of higher water vapour abun-
 538 dance (< 0.005 ppmv), where models already show good agreement with observations.

539 Heterogeneous chemistry affects the variation in hydroxyl radicals. Therefore, es-
 540 tablishing the nature of heterogeneous reactions is important in understanding the vari-
 541 ation in hydroxyl radicals, which themselves are crucial to the stability of the atmosphere
 542 e.g. McElroy and Donahue (1972); Clancy and Nair (1996). The presence of water-ice
 543 clouds could indirectly affect the recombination of carbon monoxide and molecular oxy-
 544 gen, which is catalysed by hydroxyl radicals. Future work could include establishing the

545 full spatial and temporal impact of heterogeneous chemistry on ozone and hydroxyl rad-
 546 icals by implementing the improved chemical scheme into a 3-D model, as well as inves-
 547 tigating how the change in HO_x in water-ice clouds could impact other species aside from
 548 ozone.

549 Acknowledgments

550 NOMAD vertical profiles and 1-D MPM simulations used for the plots in this study
 551 can be found in Brown et al. (2022). The full ozone dataset can be found in M. Patel
 552 et al. (2021).

553 MB was supported by the Science and Technology Facilities Council under grant
 554 number ST/P006760/1, the DISCnet Centre for Doctoral Training in Data-Intensive Sci-
 555 ence. MB would like to thank E. Brown Dewhurst for help in proof reading earlier ver-
 556 sions of this manuscript. MB would also like to thank F. Lefèvre for providing the origi-
 557 nal ASIS chemical scheme used in the 1-D model, and the Laboratoire de Météorologie
 558 Dynamique for their continuous contribution and collaboration with the GCM. SL was
 559 support by the UK Space Agency under grant numbers ST/W002949/1, ST/V005332/1.
 560 AB was support by the Engineering and Physical Sciences Research Council under grant
 561 numbers EP/V026747/1, EP/R013144/1.

562 The NOMAD experiment is led by the Royal Belgian Institute for Space Aeron-
 563 omy (IASB-BIRA), assisted by Co-PI teams from Spain (IAA-CSIC), Italy (INAF-IAPS),
 564 and the United Kingdom (Open University). This project acknowledges funding by the
 565 Belgian Science Policy Office (BELSPO), with the financial and contractual coordina-
 566 tion by the UK Space Agency through grants ST/V002295/1, ST/V005332/1 and ST/S00145X/1.
 567 This work was supported by the Belgian Fonds de la Recherche Scientifique – FNRS un-
 568 der grant number 30442502 (ET_HOME). The IAA/CSIC team acknowledges financial
 569 support from the State Agency for Research of the Spanish MCIU through the ‘Center
 570 of Excellence Severo Ochoa’ award for the Instituto de Astrofísica de Andalucía (SEV-
 571 2017-0709). US investigators were supported by the National Aeronautics and Space Ad-
 572 ministration.

573 References

- 574 Anbar, A., Leu, M.-T., Nair, H., & Yung, Y. (1993). Adsorption of HO_x on aerosol
 575 surfaces: Implications for the atmosphere of mars. *Journal of Geophysical Re-*
 576 *search: Planets*, 98(E6), 10933–10940.
- 577 Arattano, M., & Marchi, L. (2005). Measurements of debris flow velocity through
 578 cross-correlation of instrumentation data. *Natural Hazards and Earth System*
 579 *Sciences*, 5(1), 137–142. Retrieved from [https://nhess.copernicus.org/](https://nhess.copernicus.org/articles/5/137/2005/)
 580 [articles/5/137/2005/](https://nhess.copernicus.org/articles/5/137/2005/) doi: 10.5194/nhess-5-137-2005
- 581 Barth, C. A., Hord, C. W., Stewart, A. I., Lane, A. L., Dick, M. L., & Anderson,
 582 G. P. (1973). Mariner 9 ultraviolet spectrometer experiment: Seasonal varia-
 583 tion of ozone on mars [Journal Article]. *Science*, 179(4075), 795–796. Retrieved
 584 from [http://science.sciencemag.org/content/sci/179/4075/795.full](http://science.sciencemag.org/content/sci/179/4075/795.full.pdf)
 585 [.pdf](http://science.sciencemag.org/content/179/4075/795.long)<http://science.sciencemag.org/content/179/4075/795.long> doi:
 586 10.1126/science.179.4075.795
- 587 Bell III, J. F., Wolff, M. J., Malin, M. C., Calvin, W. M., Cantor, B. A., Caplinger,
 588 M. A., ... Thomas, P. C. (2009). Mars reconnaissance orbiter mars color
 589 imager (marci): Instrument description, calibration, and performance. *Jour-*
 590 *nal of Geophysical Research: Planets*, 114(E8). Retrieved from [https://](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2008JE003315)
 591 agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2008JE003315 doi:
 592 10.1029/2008JE003315
- 593 Benson, J. L., Kass, D. M., & Kleinböhl, A. (2011). Mars’ north polar hood as

- 594 observed by the mars climate sounder [Journal Article]. *Journal of Geo-*
 595 *physical Research: Planets*, 116(E3). Retrieved from [https://agupubs](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2010JE003693)
 596 [.onlinelibrary.wiley.com/doi/abs/10.1029/2010JE003693](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2010JE003693)[https://](https://agupubs.onlinelibrary.wiley.com/doi/pdf/10.1029/2010JE003693)
 597 agupubs.onlinelibrary.wiley.com/doi/pdf/10.1029/2010JE003693 doi:
 598 doi:10.1029/2010JE003693
- 599 Benson, J. L., Kass, D. M., Kleinböhl, A., McCleese, D. J., Schofield, J. T., & Tay-
 600 lor, F. W. (2010). Mars' south polar hood as observed by the mars climate
 601 sounder [Journal Article]. *Journal of Geophysical Research: Planets*, 115(E12).
 602 Retrieved from [https://agupubs.onlinelibrary.wiley.com/doi/abs/](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2009JE003554)
 603 [10.1029/2009JE003554](https://agupubs.onlinelibrary.wiley.com/doi/pdf/10.1029/2009JE003554)[https://agupubs.onlinelibrary.wiley.com/doi/](https://agupubs.onlinelibrary.wiley.com/doi/pdf/10.1029/2009JE003554)
 604 [pdf/10.1029/2009JE003554](https://agupubs.onlinelibrary.wiley.com/doi/pdf/10.1029/2009JE003554) doi: doi:10.1029/2009JE003554
- 605 Bertaux, J.-L., Korabiev, O., Perrier, S., Quémerais, E., Montmessin, F., Leblanc,
 606 F., ... Guibert, S. (2006). Spicam on mars express: Observing modes
 607 and overview of uv spectrometer data and scientific results [Journal Ar-
 608 ticle]. *Journal of Geophysical Research: Planets*, 111(E10). Retrieved
 609 from [https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2006JE002690)
 610 [2006JE002690](https://agupubs.onlinelibrary.wiley.com/doi/pdf/10.1029/2006JE002690)[https://agupubs.onlinelibrary.wiley.com/doi/pdf/](https://agupubs.onlinelibrary.wiley.com/doi/pdf/10.1029/2006JE002690)
 611 [10.1029/2006JE002690](https://agupubs.onlinelibrary.wiley.com/doi/pdf/10.1029/2006JE002690) doi: doi:10.1029/2006JE002690
- 612 Brown, M. A. J., Holmes, J. A., Sellers, G., Liuzzi, G., Lewis, S., & Patel, M. (2022,
 613 4). Dataset: Impacts of heterogeneous chemistry on vertical profiles of mar-
 614 tian ozone. Retrieved from [https://ordo.open.ac.uk/articles/dataset/](https://ordo.open.ac.uk/articles/dataset/Dataset_Impacts_of_Heterogeneous_Chemistry_on_Vertical_Profiles_of_Martian_Ozone/19646784)
 615 [Dataset_Impacts_of_Heterogeneous_Chemistry_on_Vertical_Profiles_of](https://ordo.open.ac.uk/articles/dataset/Dataset_Impacts_of_Heterogeneous_Chemistry_on_Vertical_Profiles_of_Martian_Ozone/19646784)
 616 [_Martian_Ozone/19646784](https://ordo.open.ac.uk/articles/dataset/Dataset_Impacts_of_Heterogeneous_Chemistry_on_Vertical_Profiles_of_Martian_Ozone/19646784) doi: 10.21954/ou.rd.19646784.v1
- 617 Cariolle, D., Moinat, P., Teyssèdre, H., Giraud, L., Josse, B., & Lefèvre, F.
 618 (2017). Asis v1.0: an adaptive solver for the simulation of atmospheric
 619 chemistry. *Geoscientific Model Development*, 10(4), 1467–1485. Retrieved
 620 from <https://gmd.copernicus.org/articles/10/1467/2017/> doi:
 621 10.5194/gmd-10-1467-2017
- 622 Clancy, R. T., & Nair, H. (1996). Annual (perihelion-aphelion) cycles in the pho-
 623 tochemical behavior of the global mars atmosphere [Journal Article]. *Jour-*
 624 *nal of Geophysical Research: Planets*, 101(E5), 12785-12790. Retrieved
 625 from [https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/96JE00836)
 626 [96JE00836](https://agupubs.onlinelibrary.wiley.com/doi/pdf/10.1029/96JE00836)[https://agupubs.onlinelibrary.wiley.com/doi/pdf/10.1029/](https://agupubs.onlinelibrary.wiley.com/doi/pdf/10.1029/96JE00836)
 627 [96JE00836](https://agupubs.onlinelibrary.wiley.com/doi/pdf/10.1029/96JE00836) doi: doi:10.1029/96JE00836
- 628 Clancy, R. T., Wolff, M. J., Lefèvre, F., Cantor, B. A., Malin, M. C., & Smith,
 629 M. D. (2016). Daily global mapping of mars ozone column abundances with
 630 marci uv band imaging [Journal Article]. *Icarus*, 266, 112-133. Retrieved from
 631 <http://www.sciencedirect.com/science/article/pii/S0019103515005266>
 632 doi: <https://doi.org/10.1016/j.icarus.2015.11.016>
- 633 Daerden, F., Neary, L., Viscardy, S., Muñoz, A. G., Clancy, R., Smith, M., ... Fe-
 634 dorova, A. (2019). Mars atmospheric chemistry simulations with the gem-
 635 mars general circulation model. *Icarus*, 326, 197 - 224. Retrieved from
 636 <http://www.sciencedirect.com/science/article/pii/S0019103518304883>
 637 doi: <https://doi.org/10.1016/j.icarus.2019.02.030>
- 638 Forget, F., Hourdin, F., Fournier, R., Hourdin, C., Talagrand, O., Collins, M.,
 639 ... Huot, J.-P. (1999). Improved general circulation models of the mar-
 640 tian atmosphere from the surface to above 80 km. *Journal of Geophysi-*
 641 *cal Research: Planets*, 104(E10), 24155-24175. Retrieved from [https://](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/1999JE001025)
 642 agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/1999JE001025 doi:
 643 <https://doi.org/10.1029/1999JE001025>
- 644 Holmes, J., Lewis, S., Patel, M., Chaffin, M., Cangi, E., Deighan, J., ... Van-
 645 daele, A. (2021). Enhanced water loss from the martian atmosphere dur-
 646 ing a regional-scale dust storm and implications for long-term water loss.
 647 *Earth and Planetary Science Letters*, 571, 117109. Retrieved from [https://](https://www.sciencedirect.com/science/article/pii/S0012821X21003642)
 648 www.sciencedirect.com/science/article/pii/S0012821X21003642 doi:

- 649 <https://doi.org/10.1016/j.epsl.2021.117109>
- 650 Holmes, J. A., Lewis, S. R., & Patel, M. R. (2020). Openmars: A global record
651 of martian weather from 1999 to 2015. *Planetary and Space Science*, 188,
652 104962. Retrieved from [https://www.sciencedirect.com/science/article/
653 pii/S0032063319303617](https://www.sciencedirect.com/science/article/pii/S0032063319303617) doi: <https://doi.org/10.1016/j.pss.2020.104962>
- 654 Holmes, J. A., Lewis, S. R., Patel, M. R., & Lefèvre, F. (2018). A reanalysis
655 of ozone on mars from assimilation of spicam observations [Journal Arti-
656 cle]. *Icarus*, 302, 308-318. Retrieved from [http://www.sciencedirect
657 .com/science/article/pii/S0019103517302889](http://www.sciencedirect.com/science/article/pii/S0019103517302889)[https://ac.els-cdn
658 .com/S0019103517302889/1-s2.0-S0019103517302889-main.pdf?
659 _tid=22af08de-4263-48e8-8b2a-ec0c20c3647e&acdnat=1539352983
660 _3816efbb662a3bb2f63f199e604ffdd9](https://ac.els-cdn.com/S0019103517302889/1-s2.0-S0019103517302889-main.pdf?_tid=22af08de-4263-48e8-8b2a-ec0c20c3647e&acdnat=1539352983_3816efbb662a3bb2f63f199e604ffdd9) doi: [https://doi.org/10.1016/
661 j.icarus.2017.11.026](https://doi.org/10.1016/j.icarus.2017.11.026)
- 662 Khayat, A. S. J., Smith, M. D., Wolff, M., Daerden, F., Neary, L., Patel, M. R.,
663 ... López-Moreno, J. J. (2021). Exomars tgo/nomad-uvis vertical profiles of
664 ozone: 2. the high-altitude layers of atmospheric ozone. *Journal of Geophys-
665 ical Research: Planets*, 126(11), e2021JE006834. Retrieved from [https://
666 agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2021JE006834
667 \(e2021JE006834 2021JE006834\)](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2021JE006834) doi: <https://doi.org/10.1029/2021JE006834>
- 668 Langmuir, I. (1918). The adsorption of gases on plane surfaces of glass, mica
669 and platinum. *Journal of the American Chemical Society*, 40(9), 1361-
670 1403. Retrieved from <https://doi.org/10.1021/ja02242a004> doi:
671 10.1021/ja02242a004
- 672 Lefèvre, F., Bertaux, J.-L., Clancy, R. T., Encrenaz, T., Fast, K., Forget, F., ...
673 Perrier, S. (2008). Heterogeneous chemistry in the atmosphere of mars
674 [Journal Article]. *Nature*, 454, 971. Retrieved from [http://dx.doi.org/
675 10.1038/nature07116](http://dx.doi.org/10.1038/nature07116)<http://www.nature.com/articles/nature07116>
676 doi: 10.1038/nature07116[https://www.nature.com/articles/nature07116#
677 supplementary-information](https://www.nature.com/articles/nature07116#supplementary-information)
- 678 Lefèvre, F., Lebonnois, S., Montmessin, F., & Forget, F. (2004). Three-
679 dimensional modeling of ozone on mars [Journal Article]. *Journal of Geo-
680 physical Research: Planets*, 109(E7). Retrieved from [https://agupubs
681 .onlinelibrary.wiley.com/doi/abs/10.1029/2004JE002268](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2004JE002268)[https://
682 agupubs.onlinelibrary.wiley.com/doi/pdf/10.1029/2004JE002268](https://agupubs.onlinelibrary.wiley.com/doi/pdf/10.1029/2004JE002268) doi:
683 doi:10.1029/2004JE002268
- 684 Lefèvre, F., Trokhimovskiy, A., Fedorova, A., Baggio, L., Lacombe, G., Määttänen,
685 A., ... Montmessin, F. (2021). Relationship between the ozone and water va-
686 por columns on mars as observed by spicam and calculated by a global climate
687 model. *Journal of Geophysical Research: Planets*, 126(4), e2021JE006838.
688 Retrieved from [https://agupubs.onlinelibrary.wiley.com/doi/abs/
689 10.1029/2021JE006838](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2021JE006838) (e2021JE006838 2021JE006838) doi: [https://doi.org/
690 10.1029/2021JE006838](https://doi.org/10.1029/2021JE006838)
- 691 Liuzzi, G., Villanueva, G. L., Crismani, M. M., Smith, M. D., Mumma, M. J., Daer-
692 den, F., ... Patel, M. R. (2020). Strong variability of martian water ice clouds
693 during dust storms revealed from exomars trace gas orbiter/nomad [Journal
694 Article]. *Journal of Geophysical Research: Planets*, 125(4), e2019JE006250.
695 Retrieved from [https://agupubs.onlinelibrary.wiley.com/doi/abs/
696 10.1029/2019JE006250](https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2019JE006250) (e2019JE006250 2019JE006250) doi: 10.1029/
697 2019JE006250
- 698 McElroy, M. B., & Donahue, T. M. (1972). Stability of the martian atmosphere.
699 *Science*, 177(4053), 986-988. Retrieved from [http://www.jstor.org/stable/
700 1734144](http://www.jstor.org/stable/1734144)
- 701 Montmessin, F., & Lefèvre, F. (2013). Transport-driven formation of a polar
702 ozone layer on mars [Journal Article]. *Nature Geoscience*, 6, 930. Retrieved
703 from <http://dx.doi.org/10.1038/ngeo1957><http://www.nature.com/>

- 704 articles/ngeo1957 doi: 10.1038/ngeo1957<https://www.nature.com/articles/ngeo1957#supplementary-information>
- 705
- 706 Nair, H., Allen, M., Anbar, A. D., Yung, Y. L., & Clancy, R. (1994). A photochemical model of the martian atmosphere. *Icarus*, *111*(1), 124 - 150.
- 707 Retrieved from <http://www.sciencedirect.com/science/article/pii/S0019103584711377> doi: <https://doi.org/10.1006/icar.1994.1137>
- 708
- 709 Patel, M., Mason, J., Sellers, G., Holmes, J., & Streeter, P. (2021, 10). NOMAD-UVIS ozone and aerosol vertical profile retrievals for Mars Year 34-35. Retrieved from https://ordo.open.ac.uk/articles/dataset/NOMAD-UVIS_ozone_vertical_profile_retrievals_for_Mars_Year_34-35/13580336 doi: 10.21954/ou.rd.13580336.v2
- 710
- 711
- 712
- 713
- 714
- 715 Patel, M. R., Antoine, P., Mason, J., Leese, M., Hathi, B., Stevens, A. H., ... Lopez-Moreno, J. J. (2017). Nomad spectrometer on the exomars trace gas orbiter mission: part 2—design, manufacturing, and testing of the ultraviolet and visible channel [Journal Article]. *Applied Optics*, *56*(10), 2771-2782. Retrieved from <http://ao.osa.org/abstract.cfm?URI=ao-56-10-2771><https://www.osapublishing.org/ao/abstract.cfm?uri=ao-56-10-2771> doi: 10.1364/AO.56.002771
- 716
- 717
- 718
- 719
- 720
- 721
- 722 Patel, M. R., Sellers, G., Mason, J. P., Holmes, J. A., Brown, M. A. J., Lewis, S. R., ... Lopez-Moreno, J.-J. (2021). Exomars tgo/nomad-uvis vertical profiles of ozone: 1. seasonal variation and comparison to water. *Journal of Geophysical Research: Planets*, *126*(11), e2021JE006837. Retrieved from <https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2021JE006837> (e2021JE006837 2021JE006837) doi: <https://doi.org/10.1029/2021JE006837>
- 723
- 724
- 725
- 726
- 727
- 728 Peppas, M. V., Mills, J. P., Moore, P., Miller, P. E., & Chambers, J. E. (2017). Brief communication: Landslide motion from cross correlation of uav-derived morphological attributes. *Natural Hazards and Earth System Sciences*, *17*(12), 2143–2150. Retrieved from <https://nhess.copernicus.org/articles/17/2143/2017/> doi: 10.5194/nhess-17-2143-2017
- 729
- 730
- 731
- 732
- 733 Perrier, S., Bertaux, J. L., Lefèvre, F., Lebonnois, S., Korabely, O., Fedorova, A., & Montmessin, F. (2006). Global distribution of total ozone on mars from spicam/mex uv measurements [Journal Article]. *Journal of Geophysical Research: Planets*, *111*(E9). Retrieved from <https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2006JE002681> doi: doi:10.1029/2006JE002681
- 734
- 735
- 736
- 737
- 738 Pouvesle, N., Kippenberger, M., Schuster, G., & Crowley, J. N. (2010). The interaction of h2o2 with ice surfaces between 203 and 233 k. *Phys. Chem. Chem. Phys.*, *12*, 15544-15550. Retrieved from <http://dx.doi.org/10.1039/C0CP01656J> doi: 10.1039/C0CP01656J
- 739
- 740
- 741
- 742 Shimazaki, T., & Shimizu, M. (1979). The seasonal variation of ozone density in the martian atmosphere [Journal Article]. *Journal of Geophysical Research: Space Physics*, *84*(A4), 1269-1276. Retrieved from <https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/JA084iA04p01269><https://agupubs.onlinelibrary.wiley.com/doi/pdf/10.1029/JA084iA04p01269> doi: doi:10.1029/JA084iA04p01269
- 743
- 744
- 745
- 746
- 747
- 748 Smith, M. D. (2004). Interannual variability in tes atmospheric observations of mars during 1999–2003 [Journal Article]. *Icarus*, *167*(1), 148-165. Retrieved from <http://www.sciencedirect.com/science/article/pii/S0019103503002872>https://ac.els-cdn.com/S0019103503002872/1-s2.0-S0019103503002872-main.pdf?_tid=134d7518-6e86-4870-946c-888afd3e1f3a&acdnat=1543573383_b8fcadfffcc7c60134deee13bb5416dd doi: <https://doi.org/10.1016/j.icarus.2003.09.010>
- 749
- 750
- 751
- 752
- 753
- 754
- 755 Steele, L. J., Lewis, S. R., Patel, M. R., Montmessin, F., Forget, F., & Smith, M. D. (2014). The seasonal cycle of water vapour on mars from assimilation of thermal emission spectrometer data. *Icarus*, *237*, 97–115.
- 756
- 757
- 758 Vandaele, A. C., Lopez-Moreno, J.-J., Patel, M. R., Bellucci, G., Daerden, F., Ristic,

- 759 B., . . . Team, t. N. (2018). Nomad, an integrated suite of three spectrometers
760 for the exomars trace gas mission: Technical description, science objectives
761 and expected performance [Journal Article]. *Space Science Reviews*, 214(5),
762 80. Retrieved from <https://doi.org/10.1007/s11214-018-0517-2><https://link.springer.com/content/pdf/10.1007%2Fs11214-018-0517-2.pdf> doi:
763 10.1007/s11214-018-0517-2
764
- 765 Villanueva, G. L., G., L., Aoki, S., Stone, S. W., Brines, A., Thomas, I. R., . . . Van-
766 daele, A. C. (2022). The deuterium isotopic ratio of water released from the
767 martian caps as measured with tgo/nomad. *Geophysical Research Letters*.
768 (Under Review)
- 769 Viúdez-Moreiras, D., Saiz-Lopez, A., Blaszcak-Boxe, C., Manfredi, J. R., & Yung,
770 Y. (2019). Diurnal variation in mars equatorial odd oxygen species: Chemical
771 production and loss mechanisms. *Icarus*, 113458.
- 772 Wolff, M. J., Clancy, R. T., Kahre, M. A., Haberle, R. M., Forget, F., Cantor, B. A.,
773 & Malin, M. C. (2019). Mapping water ice clouds on mars with mro/marci
774 [Journal Article]. *Icarus*. Retrieved from <http://www.sciencedirect.com/science/article/pii/S0019103518307176> doi: <https://doi.org/10.1016/j.icarus.2019.05.041>
775
776