

19 **Abstract**

20 The effects of precipitation (Pr) and evapotranspiration (ET) on soil moisture play an essential
21 role in the land-atmosphere system. Here we evaluate multimodel differences of these effects
22 within the Coupled Model Intercomparison Project Phase 5 (CMIP5) compared to Soil Moisture
23 Active Passive (SMAP) products in the frequency domain. The variability of surface soil
24 moisture (SSM), Pr, and ET within three frequency bands (7 ~ 30 days, 30 ~ 90 days, and 90 ~
25 365 days) after normalization is quantified using Fourier transform. We then analyze the impact
26 of ET and Pr on SSM variability based on a transfer function assuming these variables with a
27 linear time-invariant (LTI) system. For the simulated effects of ET and Pr on SSM variability,
28 models underestimate them in the two higher frequency bands and overestimate them in the
29 lowest frequency band but show better estimates in transitional zones between dry and wet
30 climates. Besides, the effects on SSM by Pr and ET are found to be different across the three
31 frequency bands, and models underestimate the one of Pr and ET as the dominant factor
32 controlling SSM variability in each frequency band. This study identifies the spatiotemporal
33 distribution of the CMIP5 model deficiencies in simulating ET and Pr effects on SSM.
34 Overcoming these deficiencies could improve the interpretability and predictability of Earth
35 system models in simulating interactions among the three variables.

36 **Plain Language Summary**

37 Surface climate is influenced by the interactions between the land surface and atmosphere
38 boundary, and soil moisture is a key component of these physical processes. Precipitation and
39 evapotranspiration, as two major variables involved in these interactions, have been largely
40 regarded as essential processes affecting soil moisture dynamics. However, Earth system models
41 have large uncertainties in simulating these effects. This study identifies that (1) models

42 underestimate the total effect of precipitation and evapotranspiration on soil moisture variability
43 at weekly to seasonal time scales and overestimate it at seasonal to annual time scales; (2) soil
44 moisture is mainly affected by precipitation at shorter scales and by evapotranspiration at longer
45 time scales, and models underestimate the degree of this control over the whole weekly to annual
46 frequency band; (3) model generally have better performance in the transitional climate regions
47 on capturing the effects of precipitation and evapotranspiration on soil moisture. This study
48 reveals the deficiencies of Earth system models in simulating the relationships between soil
49 moisture, precipitation, and evapotranspiration compared to satellite observations, which will
50 help improve the quantification of soil moisture dynamics in these models.

51 **1 Introduction**

52 As one of the essential components in the Earth system, soil moisture plays an important
53 role in land-atmosphere interactions (Green et al., 2019; Koster et al., 2004; Seneviratne et al.,
54 2006; Seneviratne et al., 2010). The exploration and quantification of land-atmosphere
55 interactions are significant for Earth system study and climate-change projections (Santanello et
56 al., 2018; Seneviratne et al., 2010; Suni et al., 2015).

57 The dynamics of soil moisture (SM) depend on the interplay between variability in
58 multiple hydrological processes, such as precipitation, interception, evapotranspiration, runoff,
59 and drainage (Bonan, 1996). Since these processes are complex and show large heterogeneity
60 spatiotemporally, it is hard to quantify the effects of their resulting impact on soil moisture. We
61 here focus on the two largest fluxes: precipitation (Pr), which is the water source of soil moisture
62 and also one of the atmospheric forcing variables for land surface processes; and
63 evapotranspiration (ET), which is a primary water loss relative to soil moisture.

64 Both soil moisture-precipitation (SM-Pr) and soil moisture-evapotranspiration (SM-ET)
65 interactions are some of the central issues in the climate research community and have been
66 studied for a while (Berg and Sheffield, 2018; Dong et al., 2020; Koster et al., 2004; Seneviratne
67 et al., 2010; Wang et al., 2007; Wei and Dirmeyer, 2012). Basically, soil moisture-atmosphere
68 coupling can be separated into two parts: the coupling between SM and ET and the coupling
69 between ET and Pr (Guo et al., 2016; Seneviratne et al., 2010; Wei and Dirmeyer, 2010). The
70 SM-ET coupling is linked to the impact of SM on ET variability as a regulator of energy
71 partitioning (Seneviratne et al., 2010) and is mostly a local process (Wei and Dirmeyer, 2012).
72 On the other hand, the SM-Pr coupling, which includes the effect of SM on ET and the effect of
73 ET on Pr, is more elusive due to the series of atmospheric processes, especially the interactions
74 between ET and Pr (see Seneviratne et al., 2010).

75 Studies on the effects of Pr and ET on the temporal variability of SM focused on
76 analyzing autocorrelations of SM time series. Considering the SM dynamics as being forced by a
77 random precipitation time series (i.e., white noise) and damped by an exponential damping term
78 related to evapotranspiration losses, the temporal variability of SM can be reasonably governed
79 by a first-order Markov process, which results in the SM time series to exhibiting a red noise
80 spectrum (Delworth and Manabe, 1988). Based on this, many studies worked on characterizing
81 these effects from a time-frequency domain. The response of SM to Pr at long time scales was
82 investigated and revealed the amplitude decrease and the phase shift of soil hydrology with soil
83 depth (Wu et al., 2002). This phase shift as to how SM spectra related to Pr was further explored
84 using the integral time scale to show that SM spectra decay more rapidly than a red noise due to
85 Pr departing from white noise at high frequency, and the damping term of ET losses was found
86 to be bounded by the maximum of ET (Katul et al., 2007). Similarly, the integral time scale was

87 used to reveal the dynamics of SM memory and its correlation with Pr and ET (Ghannam et al.,
88 2016). Based on previous studies (Katul et al., 2007), the SM spectrum could not be explained
89 only by precipitation effects on longer time scales (Nakai et al., 2014). This concept has also
90 been used to investigate the effects of Pr on SM variability on a regional scale (Zhou et al.,
91 2020).

92 Although the SM-Pr and SM-ET couplings have been studied for a long time, the effects
93 of Pr and ET on soil moisture variability are still not completely understood (Seneviratne et al.,
94 2010). Even less understood is how Earth system models perform in capturing these effects
95 globally and at different time scales. There are two major challenges. One is the lack of enough
96 in-situ soil moisture measurements at the global scale. Nowadays, remote sensing technology,
97 such as NASA's Soil Moisture Active Passive (SMAP) mission (Entekhabi et al., 2010), provides
98 global observation of soil moisture at a high spatiotemporal resolution that can be used to
99 constrain land-atmosphere interaction observations over different spatiotemporal scales (Guilod
100 et al., 2015; Tuttle and Salvucci, 2016). Additionally, although it only provides surface soil
101 moisture (top ~5cm of the soil column), several studies have shown that surface soil moisture
102 (SSM) and root-zone soil moisture (RZSM) have strong correlations in quantifying surface flux
103 (Akbar et al., 2018; Ford et al., 2014; Qiu et al., 2016), indicating that SSM can be regarded as a
104 proxy for RZSM under most conditions (McColl et al., 2019). Another challenge is that, due to
105 the complexity and the large number of processes involved in land-atmosphere interactions, the
106 representation of couplings between SM, Pr, and ET highly relies on parameterizations within
107 Earth system models, which leads to large uncertainties in identifying the effects of Pr and ET on
108 SM variability. The transfer function (Haykin and Van Veen, 2007), as a mathematical
109 representation of the differential equation of system dynamics, can be used to describe the

110 relationship between the signal input and response assuming a linear time-invariant (LTI) system
111 (Phillips et al., 2003) using a time-frequency analysis, without considering its specific structure
112 and parameters. Therefore, it can be used to investigate the effects of Pr and ET on SM in the
113 frequency domain, assuming they are nearly an LTI system. The spectral analysis based on the
114 LTI system has been applied to other hydrological research like the runoff-storage relationship
115 (Riegger and Tourian 2014) and the surface flow in the river during floods (Bailly-Comte et al.,
116 2008).

117 The fifth phase of the Coupled Model Intercomparison Project (CMIP5; Taylor et al.
118 2012), which integrated a set of model experiments to improve our knowledge of climate change
119 and climate variability, provides an opportunity for the multimodel assessment of land-
120 atmospheric processes and variability. Evaluation of CMIP5 has been the ongoing interest of the
121 research community (Yuan et al., 2021). Although evaluations of land-atmosphere interactions
122 related to soil moisture within CMIP5 have been performed earlier (e.g., Berg and Sheffield,
123 2018; Dirmeyer et al., 2013; Levine et al., 2016), few studies have characterized the temporal
124 behavior of SM globally in order to illustrate the model performance across frequency regimes.
125 Therefore, this study takes advantage of the CMIP5 intercomparison project to evaluate 14 Earth
126 system models (ESMs) in simulating the effects of Pr and ET on SSM variability. We aim to
127 address two main objectives in this study: 1) how the effects of Pr and ET on SSM variability are
128 at different time scales, and 2) how the ESMs within CMIP5 perform in capturing these effects.
129 Specifically, these effects are analyzed within three frequency bands: 1) weekly to monthly time
130 scales ($1/7 \sim 1/30 \text{ day}^{-1}$), 2) monthly to seasonal time scales ($1/30 \sim 1/90 \text{ day}^{-1}$), and 3) seasonal
131 to annual time scales ($1/90 \sim 1/365 \text{ day}^{-1}$) at the global scale. Further, a Fourier analysis is
132 conducted to determine the variability and power spectra over those various periods (Thomson

133 and Emery, 2014; Wilks, 2011). Similar approaches to decomposing the time series into different
134 frequency bands have been used to understand the precipitation and soil moisture variability
135 (Ruane and Roads, 2007; Wei et al., 2010; Xi et al., 2022).

136 In section 2, we first describe the models and data used. Then, we detail our methodology
137 for spectral analysis. In section 3, we show the results of observation-based data in the first part.
138 In the second part, we perform comparative analyses to evaluate the multimodel differences
139 within CMIP5. In the third part, we investigate uncertainties that may exist in this study. Finally,
140 in section 4, we summarize our findings and discuss the impacts of the research.

141 **2 Methods**

142 2.1 Overview

143 We first describe the data collection within CMIP5, SMAP observation data, and ERA5
144 reanalysis data. Second, we detail the methodology from data preprocessing to the final
145 multimodel comparison (Figure 1). Specifically, section 2.3 describes the preprocessing of
146 SMAP products and CMIP5 simulations. Section 2.4 defines the normalized variability of SSM,
147 Pr, and ET and how to get them within the three frequency bands. Next, section 2.5 introduces
148 two ratios used to investigate the effects of Pr and ET on SSM based on the normalized
149 variability defined in section 2.4. Section 2.6 describes the spectral slopes of SSM, Pr, and ET
150 time series and how to depict them as the color of noise. Finally, section 2.7 describes how the
151 models are compared with the observation-based data and illustrates the significance test.

152 2.2 Data Organizing

153 CMIP5 integrated a set of model experiments to improve our knowledge of climate
154 variability from past to present to future (Taylor et al., 2012). Here we use the daily simulations

155 of 14 ESMs from the historical experiment within CMIP5. The models are selected based on the
156 availability of daily outputs required for the spectral analysis within the same temporal coverage
157 from 01/01/1950 to 12/31/2005 (Table S1). To evaluate the effects of Pr and ET (i.e.,
158 atmospheric water supply and loss) on SSM variability, we analyze the simulated SSM (top 10
159 cm), Pr, and ET (variable *mrsos*, *pr*, and *hfls* in the CMIP5 archive, respectively). We use only
160 one ensemble member – “r1i1p1” (where *r* for realization, *i* for initialization, and *p* for physics).

161 Observation data of SSM are taken from SMAP (Entekhabi et al., 2010). For Pr and ET,
162 we use reanalysis data from ERA5 (Copernicus Climate Change Service (C3S), 2017), the fifth-
163 generation reanalysis of ECMWF (European Centre for Medium-Range Weather Forecasts) as
164 the next generation of representative satellite-observational data, as a reference to compare with
165 CMIP5 simulations on the global scale. To ensure that the data are consistent, we use datasets
166 from SMAP and ERA5 with the same temporal coverage, spanning 1 April 2015 to 31 December
167 2020.

168 The NASA SMAP satellite was launched in January 2015 and has been measuring SSM
169 (moisture in the top ~5 cm of the soil column) globally every 2~3 days (Entekhabi et al., 2010).
170 SMAP soil moisture matches well *in situ* SSM observations (Chan et al., 2016, 2018; Colliander
171 et al., 2017, 2021) and shows higher accuracy measured by a global average anomaly correlation
172 over the majority of available land pixels compared to two other satellite products (Chen et al.,
173 2018). Additionally, SMAP has been shown to have high information content relative to four
174 other retrieval products of soil moisture (Kumar et al., 2018). In this study, we use its Level 3
175 Radiometer Global Daily 36 km EASE-Grid Soil Moisture, Version 7 (O’Neill et al., 2020) with
176 the retrievals from both 6 am descending passes and 6 pm ascending passes. Although its 6 pm
177 retrievals show more degradation than its 6 am retrievals due to the required vertical thermal

178 equilibrium assumption in its algorithm, this degradation was shown to be small (Chan et al.,
179 2018; O’Neill et al., 2018). Therefore, we use both retrievals to best use the observational
180 information. The Level 3 product of SMAP was developed based on geophysical parameters
181 derived from its Level 1 and Level 2 products. It was spatiotemporally re-sampled to the global
182 cylindrical EASE-Grid 2.0 to make each grid cell has a nominal size of approximately 36×36
183 km² regardless of longitude and latitude (Brodzik et al., 2012).

184 The reference observation-based data of precipitation (Pr) and evapotranspiration (ET)
185 are collected from ERA5. ERA5 reanalysis is achieved by data assimilation, which combines
186 weather forecasts with observations in an optimal way every few hours to produce the best
187 estimate of the state of the atmosphere. In this way, ERA5 combines model data and
188 observations into a globally complete and consistent dataset. ERA5 reanalysis has been
189 evaluated extensively on regional and global scales and shows great improvements over its
190 popular predecessor ERA-Interim and is a potential reference as proxies for observations for the
191 hydrological process modeling (Jiao et al., 2021; Martens et al., 2020; Rivoire et al., 2021; Tarek
192 et al., 2020). In this study, we use “total precipitation” (units: m) and “evaporation” (unit: m of
193 water equivalent) estimates on single levels as the observation-based Pr and ET, respectively
194 (Copernicus Climate Change Service (C3S), 2017). This dataset has a spatial resolution of
195 0.25°×0.25° for the atmosphere, spanning 1979 to the present, with an hourly temporal
196 resolution. We collect the ERA5 hourly data within the same period as SMAP. Then we convert
197 them into daily total precipitation and evapotranspiration (units: m) based on (Copernicus
198 Climate Change Service (C3S), 2017):

$$Pr_d = \sum_{hr=1}^{23} Pr_{hr} + Pr_{d+1\ 00UTC} \quad (1)$$

$$ET_d = \sum_{hr=1}^{23} ET_{hr} + ET_{d+1\ 00UTC} \quad (2)$$

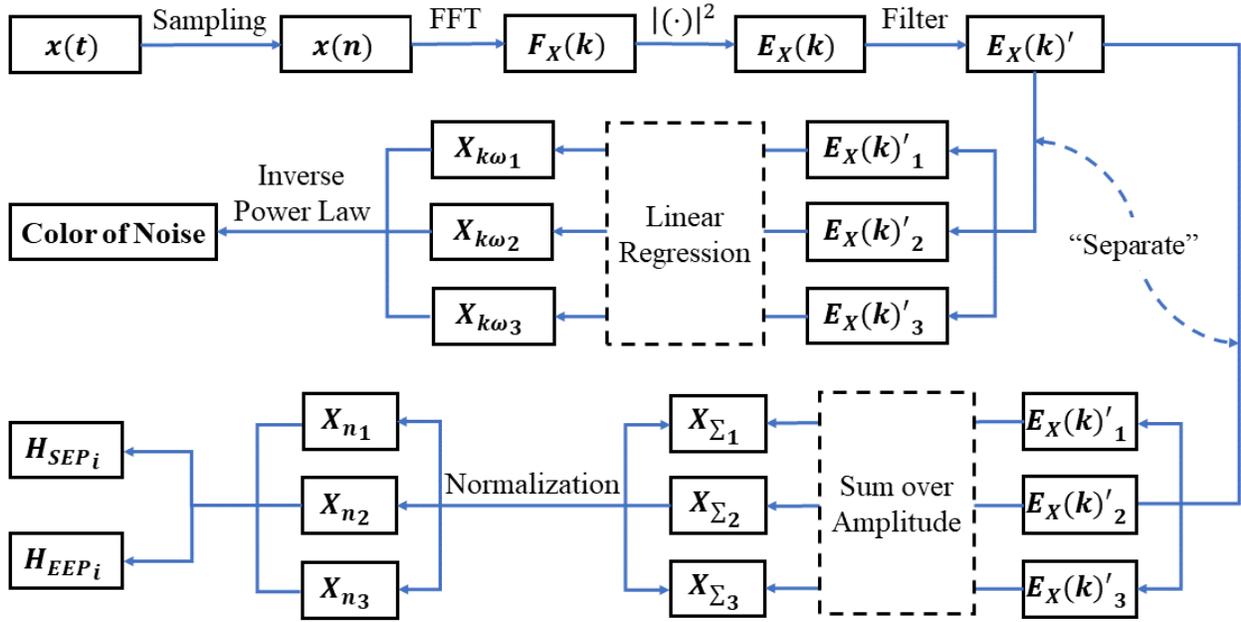
199 where hr is hour and d is the day of interest ($d + 1$ is the next day). This means that we need
 200 two days of data to get total precipitation and evapotranspiration per day. For example, to
 201 calculate total precipitation for 1 April 2015, we need hourly data on 1 April 2015 with time = 01
 202 – 23 to cover 00 – 23 UTC for 1 April 2015 and the hourly data on 2 April 2015 with time = 00
 203 to cover 23 – 24 UTC for 1 April 2015. In this way, we get daily precipitation and
 204 evapotranspiration time series (i.e., Pr_d and ET_d) for further analysis. We also use the same
 205 subset of ERA5 datasets and the same method to collect daily potential evaporation (PE) data
 206 (units: m) with the same temporal coverage for comparison with ET.

207 2.3 Data Preprocessing

208 The data in this study are preprocessed before the spectral analysis, as described in our
 209 previous study (Xi et al., 2022). Basically, since the global SSM retrievals from SMAP are
 210 temporally discontinuous on a daily time scale, we first perform a gap-filling to make it a daily
 211 dataset. For the model estimations within CMIP5 that span decades, there might be long-memory
 212 fluctuations on such time scales (Mudelsee, 2013). To avoid such long-memory trends from
 213 introducing errors into the power spectrum when performing Fourier analysis, we detrend the
 214 CMIP5 data to obtain a stationary signal by subtracting an optimal (least squares) fitted linear
 215 regression from original data. In this way, the time series after detrending has a mean value of
 216 zero, and we focus on their intra-annual fluctuations.

217 2.4 Normalized variability of SSM, Pr, and ET

218 Normalized variability of SSM (SSM_n), Pr (Pr_n), and ET (ET_n) of CMIP5 models and
 219 observation-based data are both calculated for comparison. We aim to use SSM_n , Pr_n , and ET_n to
 220 indicate the proportion of the temporal variability over different frequency bands. These
 221 normalized variabilities are further used to evaluate the effects of Pr and ET on SSM variability
 222 in section 2.5. The procedures to get SSM_n , Pr_n , and ET_n from time series of SSM ($ssm(t)$), Pr
 223 ($pr(t)$), and ET ($et(t)$), are shown in Figure 1 (for a detailed version, see Figure S1). The details
 224 of the steps to process SSM_n are explained in a previous study (Xi et al., 2022). The processing
 225 of Pr_n and ET_n follows a similar procedure. Here we give a basic idea of the strategy used to
 226 process X_n .



227
 228 **Figure 1.** Steps to get the normalized variability (X_{n1} , X_{n2} , and X_{n3} , hereafter collectively referred to as
 229 X_n) and the spectral slope ($X_{k\omega_1}$, $X_{k\omega_2}$, and $X_{k\omega_3}$, hereafter collectively referred to as X_{kw}) of the
 230 variable X from its time series ($X(t)$). X here means SSM, Pr, and ET, since the procedure to deal with
 231 $ssm(t)$, $pr(t)$, and $et(t)$ is the same. The number “1”, “2”, and “3” (hereafter being referred as i)

232 represent three frequency bands in the order of weekly to monthly (7 ~ 30 days), monthly to seasonal (30
 233 ~ 90 days), and seasonal to annual (90 ~ 365 days) time scales. $x(n)$ is the discrete series sampled from
 234 $x(t)$. $F_X(k)$ is the amplitude spectrum of X from $x(n)$ using Fast Fourier Transform (FFT). $E_X(k)$ is the
 235 power spectrum of X as the square of the absolute value of its amplitude. $E_X(k)'$ is the filtered $E_X(k)$ to a
 236 frequency band within 7 to 365 days. $E_X(k)'_i$ is $E_X(k)'$ being “separated” into the three frequency bands:
 237 weekly to monthly ($i = 1$), monthly to seasonal ($i = 2$), and seasonal to annual ($i = 3$). The sum of
 238 spectral amplitudes of X (X_{Σ_i}) and X_{kw_i} is gotten from $E_{SSM}(k)'_i$ based on “sum over amplitude” and
 239 “linear regression” within the i th frequency band, respectively. X_{n_i} is gotten from X_{Σ_i} based on
 240 normalization across the three frequency bands, and then H_{SEP_i} and H_{EEP_i} are two ratios used to analyze
 241 the effects of Pr and ET on SSM defined in section 2.5.

242 The computation of X_n for models and observations is the same. It is based on the Fast
 243 Fourier Transform (FFT), a faster algorithm for the Discrete Fourier Transform (DFT). They
 244 decompose the time series into orthogonal sinusoidal frequency components so that the
 245 variability within each component can be investigated separately. In this way, the oscillations of
 246 time series ($x(t)$) can be identified through the spectra in the frequency domain. All
 247 computations and statistical analyses in this study are programmed in MATLAB
 248 (<http://www.mathworks.com/>).

249 First, we use FFT to get the amplitude spectrum of X ($F_X(k)$) from $x(n)$, which is the
 250 discrete series sampled from $x(t)$ based on the sampling number (N) (i.e., the number of days).
 251 Then we get the power spectrum of X from its amplitude spectrum as $E_X(k) = |F_X(k)|^2$. We
 252 only keep $E_X(k)$ with the frequency ranges from $1/2$ to $1/N$ day⁻¹ since the spectrum is
 253 symmetrical about the Nyquist frequency ($f_s/2$, where f_s is sampling frequency). For all time-
 254 series data (i.e., CMIP5 simulations, SMAP, and ERA5 references), we use 1 day⁻¹ as the
 255 sampling frequency from $x(t)$ to $F_X(k)$ since they are all with daily resolution.

256 Then, we restrict our investigation within a weekly to annual frequency band by using a
 257 low-pass filter and a high-pass filter with the cutoff frequency as $1/7 \text{ day}^{-1}$ and $1/365 \text{ day}^{-1}$,
 258 respectively. Next, we separate the filtered $E_X(k)$ ($E_X(k)'$) into three frequency bands: weekly to
 259 monthly time scales (7 ~ 30 days), monthly to seasonal time scales (30 ~ 90 days), and seasonal
 260 to annual time scales (90 ~ 365 days). Finally, we define the normalized variability of X as the
 261 spectral power of each frequency band divided by the total spectral power of $E_X(k)'$:

$$X_{n_i} = \frac{\sum_j E_{X_i}(k_j)'}{\sum_{i=1}^3 \sum_j E_{X_i}(k_j)'} \quad (3)$$

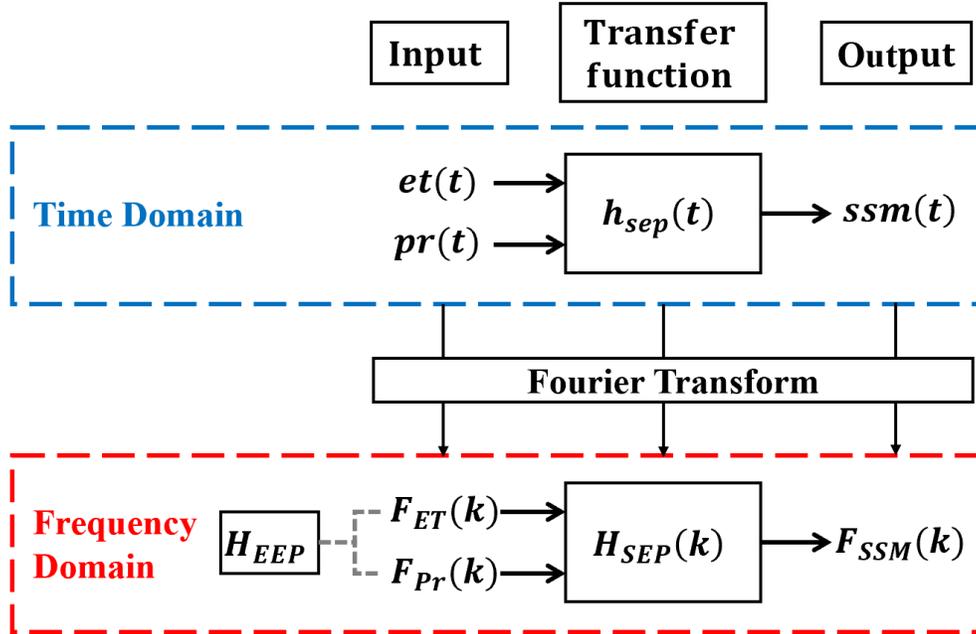
262 where $E_{X_i}(k_j)'$ represents the spectral power of X for the j th frequency in the i th frequency band,
 263 i is the ordinal number representing the three frequency bands from high to low, and j is the
 264 ordinal number of each frequency within each frequency band. Thus, we denote X_{n_i} as the
 265 normalized variability of X in the i th frequency band. In this way, X_{n_i} , as a value between 0 and
 266 1, indicates the proportion of the temporal variability of $x(t)$ in the i th frequency band.

267 2.5 Analysis of the Effects of Pr and ET on SSM Variability

268 Both Pr and ET affect SSM variability. Pr is the water source of SSM, while ET is the
 269 water loss term affecting SSM. Thus, increasing Pr will increase SSM while increasing ET will
 270 decrease SSM (without considering the saturation condition), which can be expressed as
 271 (neglecting other processes):

$$\frac{dssm(t)}{dt} = pr(t) - et(t) \quad (4)$$

272 Here we aim at examining the effects of ET and Pr on SSM variability within the three frequency
 273 bands based on the transfer function of a conceptual LTI system. The related theory of the LTI
 274 system and transfer function can be found in Appendix A.



275

276 **Figure 2.** Diagram of the conceptual LTI system with the excitations as $et(t)$ and $p(t)$, the response as
 277 $ssm(t)$, and the transfer function as $h_{sep}(t)$ in the time domain. The form in the time domain is shown in
 278 the blue box. By performing Fourier transform, the corresponding form of the LTI system in the
 279 frequency domain is shown in the red box, where $F_{ET}(k)$, $F_{Pr}(k)$, and $F_{SSM}(k)$ is the Fourier transform
 280 (amplitude spectrum) of $et(t)$, $p(t)$, and $ssm(t)$, and $H_{SEP}(k)$ is the Fourier transform of the transfer
 281 function $h_{sep}(t)$. H_{EEP} is the fraction of ET variability to the sum of ET and Pr variability in the
 282 frequency domain.

283 To capture the total effects of ET and Pr on the SSM variability, we use a conceptual LTI
 284 system with the excitation as $et(t)$ and $p(t)$ together and the response as $ssm(t)$ (Figure 2).
 285 Since ET and Pr have different spectral characteristics in the frequency domain (Katul et al.,
 286 2007; Nakai et al., 2014; also from Figure 3 in section 3.1), here we separate their effects on
 287 SSM as two inputs and determine the total effects as an identical transfer function. Regarding
 288 this system as a “black-box” model, we can focus on the relationship between excitation (i.e., ET

289 and Pr) and response (i.e., SSM) without caring about the internal variations of the system. In
 290 this way, the relationship between $ssm(t)$, $et(t)$, and $p(t)$ can be expressed :

$$ssm(t) = et(t) \otimes h_{sep}(t) + pr(t) \otimes h_{sep}(t) \quad (5)$$

291 where $h_{sep}(t)$ is the transfer function of the LTI system shown in Figure 2. Then, equation (2)
 292 can be expressed as spectrum analysis in the frequency domain:

$$F_{SSM}(k) = F_{ET}(k) \cdot H_{SEP}(k) + F_{Pr}(k) \cdot H_{SEP}(k) \quad (6)$$

293 where $H_{SEP}(k)$ is the Fourier transform of the transfer function $h_{sep}(t)$. Thus, the variations of
 294 the excitation and response spectra of the LTI system are determined by the transfer function
 295 $H_{SEP}(k)$ as:

$$H_{SEP}(k) = \frac{F_{SSM}(k)}{F_{ET}(k) + F_{Pr}(k)} \quad (7)$$

296 where $F_{ET}(k)$, $F_{Pr}(k)$, and $F_{SSM}(k)$ is the Fourier transform (amplitude spectrum) of $et(t)$,
 297 $p(t)$, and $ssm(t)$.

298 In order to characterize the total effects of ET and Pr on SSM variability within the three
 299 frequency bands, we process equation (5) based on the normalized variability (SSM_{n_i} , ET_{n_i} , and
 300 Pr_{n_i}) defined in section 2.4:

$$H_{SEP_{n_i}} = \frac{SSM_{n_i}}{ET_{n_i} + Pr_{n_i}} \quad (8)$$

301 where $H_{SEP_{n_i}}$ is the fraction of SSM variability to the sum of ET and Pr variability (i.e., demand
 302 and supply) in the i th frequency band. The higher this ratio, the stronger influences on the
 303 temporal variability of SSM by ET and Pr. We also aim to define the dominant factor on SSM
 304 variability (i.e., whether ET or Pr) within the three frequency bands. Therefore, we define
 305 another ratio:

$$H_{EEP_{n_i}} = \frac{ET_{n_i}}{ET_{n_i} + Pr_{n_i}} \quad (9)$$

306 where $H_{EEP_{n_i}}$ is the fraction of ET variability to the sum of ET and Pr variability in the i th
 307 frequency band. This ratio is greater than one-half means that ET has larger variability than Pr
 308 and thus a greater impact on the temporal variability of SSM and vice versa. In this way, we use
 309 $H_{SEP_{n_i}}$ and $H_{EEP_{n_i}}$ as two indicators to characterize the effects of ET and Pr on SSM variability
 310 in the three frequency bands – $H_{SEP_{n_i}}$ measures the total effect of ET and Pr on SSM variability
 311 and $H_{EEP_{n_i}}$ determines which process is dominant. A detailed procedure to get $H_{SEP_{n_i}}$ and
 312 $H_{EEP_{n_i}}$ can be found in Figure S1.

313 2.6 Analysis of Spectral Slope of SSM, Pr, and ET

314 The spectral slope exhibits characteristics of the soil moisture's physical behavior. This
 315 factor can explain how ET and Pr variability contribute to the spectrum of soil moisture (Katul et
 316 al., 2007). Being considered power-law noise signals, the spectral densities of time series vary as
 317 proportional to $1/f^\beta$ (i.e., inverse frequency), where β is the inverse number of the spectral
 318 slope (Bourke, 1998). In this way, the color of the noise, which is related to the power spectrum
 319 of noise signals, can be used to indicate the spectral slopes of SSM, Pr, and ET. The basic theory
 320 of the color of noise can be found in Text S2.

321 The noise colors can be divided into several types according to the slope of their power
 322 spectral density. We use white noise and five main colored noises (violet, blue, pink, red, and
 323 black noise) to characterize the spectral slopes for SSM ($SSM_{k\omega_i}$), Pr ($Pr_{k\omega_i}$), and ET ($ET_{k\omega_i}$) in
 324 the i th frequency band. The corresponding spectral slope (equal to β in inverse power law $1/f^\beta$)
 325 of violet, blue, white, pink, and red noise (or Brownian noise) is 2, 1, 0 (i.e., the spectral density

326 of white noise is flat), -1, and -2, respectively, and the spectral slope of black noise is smaller
327 than -2. The smaller the spectral slope in the frequency domain, the longer the memory of the
328 signals represented as different colors of noise (excluding violet and blue noise). For example, a
329 signal with its spectrum shown as white noise means the contribution to its variance is equal
330 across all frequencies, while a signal with its spectrum shown as red noise means low-frequency
331 periodic components dominate the contribution to its variance. Therefore, we use $SSM_{k\omega}$, $Pr_{k\omega}$,
332 and $ET_{k\omega}$ to characterize the memory of SSM, Pr, and ET. The steps to get these variables can
333 also be found in Figure 1.

334 2.7 Analysis of Differences between Models and Observational references

335 We evaluate two multimodel differences within CMIP5 compared to SMAP and ERA5
336 data: 1) differences in H_{SEP_n} and H_{EEP_n} ; and 2) differences in $SSM_{k\omega}$, $ET_{k\omega}$, and $P_{k\omega}$, by
337 subtracting observation-based data from model averages. In addition, we calculate the coefficient
338 of variation across 14 models to show the degree of the statistical dispersion of the quantities.

339 The spatial resolution and the land cover between CMIP5 models and observational
340 references (i.e., SMAP and ERA5), as well as among models themselves, are different. Here we
341 re-grid all products with the same spatial resolution (36 km×36 km) and land cover as SMAP
342 based on the nearest neighbor binning so they can be compared with each other (details on the
343 spatial resolution projection see previous work (Xi et al., 2022)). In addition, we perform a
344 significance test on these differences to avoid that multimodel differences in some regions may
345 be caused by only a few models or even one model. This significance test is depicted on the
346 maps using stippling, showing the regions that pass the 100% (i.e., all 14 models agree on the
347 sign of average differences) and 75% (i.e., 11 of the 14 models agree on the sign of average
348 differences) significance test. Since the variation of soil moisture in dry regions is usually very

349 small (Koster et al., 2009), we remove regions with \overline{SSM}_n less than 0.1 (shown as dark gray on
350 the maps), where \overline{SSM}_n is defined as the observational mean SSM after spatiotemporal
351 normalization (Figure S3).

352 **3 Results and Discussion**

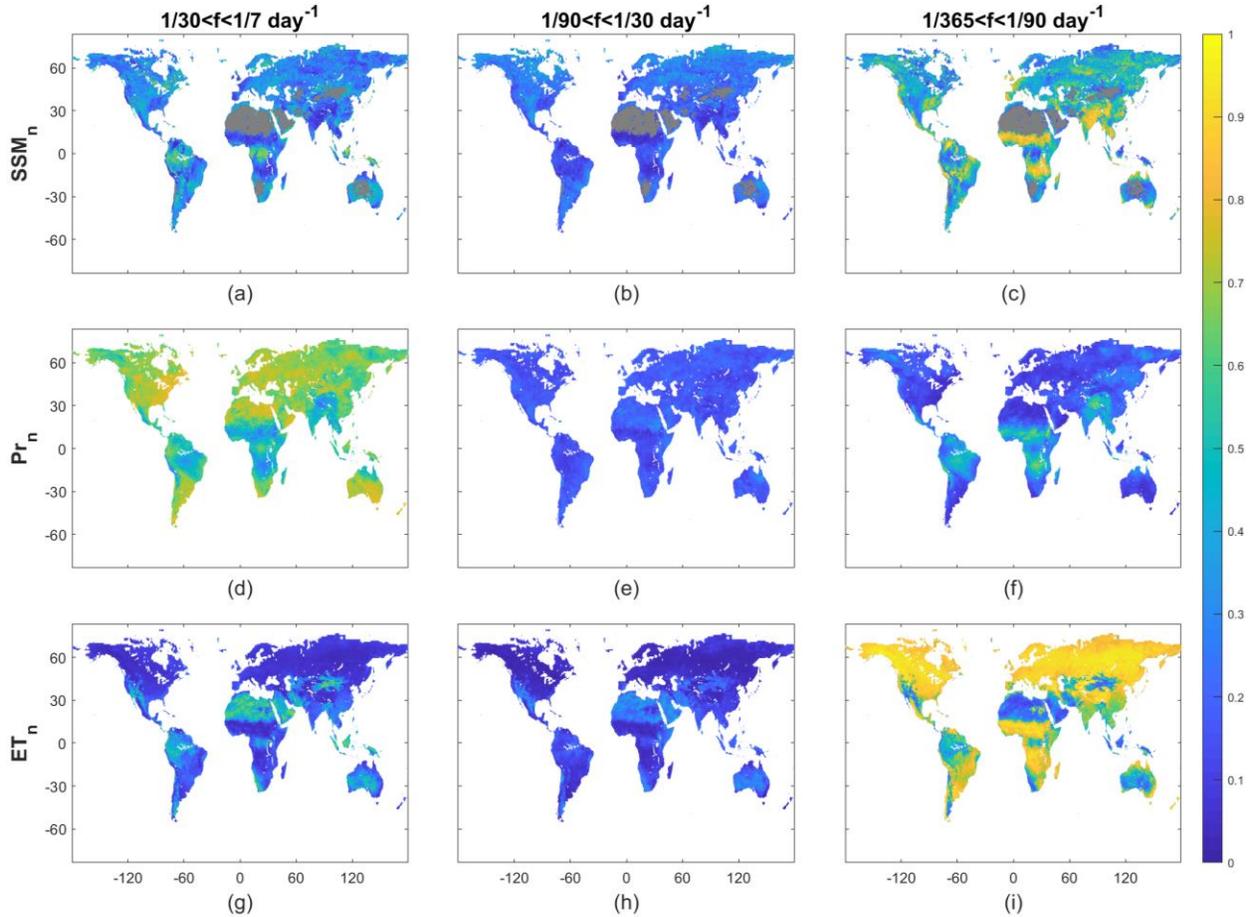
353 3.1 Temporal Variabilities of Soil Moisture, Precipitation, and Evapotranspiration from SMAP 354 and ERA5 Data

355 The temporal variability of SSM (SSM_n) concentrates more in the seasonal to annual
356 frequency band in most regions, with a smaller proportion in the two higher frequency bands,
357 indicating that SSM has a large variability on time scales longer than the seasonal time scale
358 (Figure 3a-3c).

359 The temporal variability of precipitation (Pr_n) shows different regional distributions over
360 the three frequency bands (Figure 3d-3f). The variability is larger in the lowest frequency band
361 for most tropical regions where the seasonal cycle can be large, and is larger in the highest
362 frequency band for other regions, especially non-tropical regions. The reason is that, in most
363 tropical regions, especially regions with tropical wet and dry climate, like Brazil, India, Northern
364 Australia, and regions between the Sahara Desert and the equator in Africa, although the
365 variation of temperature and radiation are small over a year, rainfall exhibits a strong seasonal
366 cycle – the days with and without rainfall are concentrated so that the boundaries of the wet
367 season and dry season are more distinct. So, precipitation in these regions shows a large seasonal
368 variability. However, in tropical regions with a very wet climate, such as the Democratic
369 Republic of the Congo, Indonesia, and the Philippines, there is no such seasonality because of
370 the more steady rainfall pattern in these regions. On the other hand, there is not an obvious wet

371 and dry season distinction for most non-tropical regions. The occurrence of rainfall is typically
372 more random over a whole year and close to a white noise signal at high frequencies (Katul et
373 al., 2007; Nakai et al., 2014). Therefore, precipitation variability in non-tropical regions is almost
374 all high-frequency variability, except for regions with a Mediterranean climate and monsoonal
375 regions where the monsoon distributes rainfall in a few months, imposing a strong seasonal
376 cycle.

377 The largest temporal variability of ET (ET_n) in the lowest frequency band means that ET
378 variability is large on time scales longer than seasonal over most regions (Figure 3g-3i), except
379 in regions with a tropical wet climate. The reason is that ET in most regions is driven by either
380 radiation or moisture limitation with high seasonality, except in the wet tropics where the
381 seasonality of radiation and moisture is small but the daily variability can be large. In this way,
382 the results in tropical wet regions, such as in the Amazon Rainforest, Africa's Equator, Indonesia,
383 and the Philippines, are the opposite of other regions in terms of frequency distribution, showing
384 ET variability concentrates on time scales shorter than monthly. This high-frequency radiation
385 variability is mainly due to the variability of clouds on daily to weekly time scales which causes
386 a large variability of ET on these short time scales (Anber et al., 2015). Moreover, this
387 mechanism has the largest influence on regions near the equator because these regions receive
388 more radiation than other regions over a year. Therefore, in these regions, ET variability is
389 mostly located in the highest frequency band. In addition, ET in very dry regions (e.g., desert)
390 does not have a clear seasonal cycle as well due to the strong limitation of moisture.

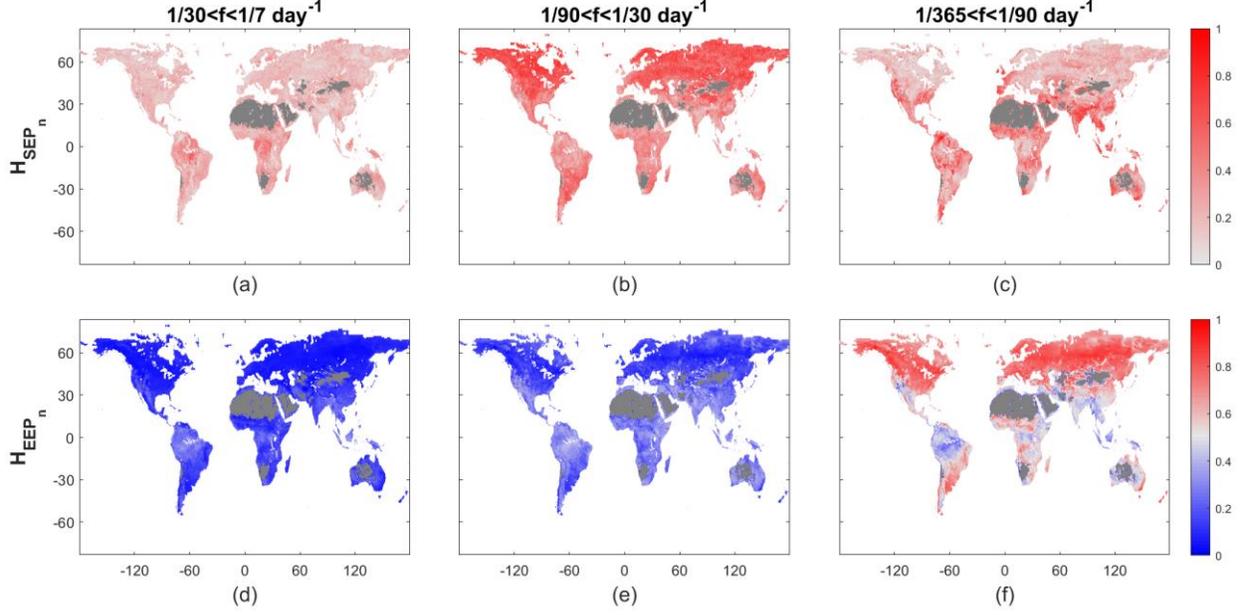


391
 392 **Figure 3.** SSM_n (Figure a-c), Pr_n (Figure d-f), and ET_n (Figure g-i) based on SMAP and ERA5 data over
 393 the three frequency bands. SSM_n , Pr_n , and ET_n is the normalized variability of SSM, Pr, and ET,
 394 respectively, defined in section 2.5. Dark grey parts in Figure a-c are regions with \overline{SSM}_n (observational
 395 mean SSM after spatiotemporal normalization) less than 0.1. For all subsequent results, including Figure
 396 3, the three columns from left to right represent the weekly to monthly frequency band ($n = 1$), the
 397 monthly to seasonal frequency band ($n = 2$), and the seasonal to annual frequency band ($n = 3$).

398 The temporal variability of Pr and ET both show an apparent regional distribution (Figure
 399 3). For Pr, the variability in tropical and non-tropical regions is opposite across the three
 400 frequency bands – the variability in tropical regions concentrates in the seasonal to annual
 401 frequency band, and the variability in non-tropical regions concentrates in the weekly to monthly

402 frequency band. For ET, the variability in most regions concentrates in the seasonal to annual
403 frequency band except for the dry regions and regions near the equator where the variability
404 concentrates in the weekly to monthly frequency band. However, compared to Pr and ET, the
405 temporal variability of SSM is more diverse spatially on a global scale.

406 Figure 4 shows the global distribution of H_{SEP_n} and H_{EEP_n} based on SMAP and ERA5
407 data over the three frequency bands (the corresponding values of H_{SEP_n} and H_{EEP_n} in each
408 frequency band see Table S4). In the weekly to monthly frequency band, the total effect of ET
409 and Pr on SSM variability is less than it in the other two frequency bands. Compared to Pr,
410 which is the dominant driver of SSM variability in this frequency band, the fluctuation of ET has
411 limited effects on SSM as ET is a slower process, in part regulated by soil moisture itself
412 (Figures 4a and 4d). On time scales longer than monthly, ET and Pr together have more effects
413 on SSM variability. In the monthly to seasonal frequency band where the total effect of ET and
414 Pr on SSM reaches its largest magnitude, although the proportion of ET variability becomes
415 larger, Pr is still the dominant factor of SSM variability (Figures 4b and 4e). In the seasonal to
416 annual frequency band, the total variability of ET and Pr decreases but is still larger than it in the
417 weekly to monthly frequency band. However, in this frequency band, ET becomes the dominant
418 factor on SSM, especially in the middle and high latitudes. Therefore, Pr variability alone in
419 these regions is no longer able to explain the SSM dynamics, and the seasonality of ET has to be
420 considered (Figures 4c and 4f). Since H_{EEP_n} represents the proportion of ET variability to the
421 total variability of ET and Pr, similar to ET_n shown in Figure 3, H_{EEP_n} patterns are different in
422 tropical wet regions, where ET variability has more effects on SSM on the two higher frequency
423 bands (Figures 4d and 4e), and Pr becomes the dominant factor on the lowest frequency band
424 due to the strong seasonality in rainfall (Figure 4f).



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Figure 4. H_{SEP_n} (Figure a-c) and H_{EEP_n} (Figure d-f) based on SMAP and ERA5 data over the three frequency bands. H_{SEP_n} is the ratio of SSM_n to the sum of ET_n and Pr_n , and H_{EEP_n} is the ratio of ET_n to the sum of ET_n and Pr_n , defined in section 2.6. The values within each frequency band are normalized to between zero and one across the three frequency bands. Dark grey parts are regions with \overline{SSM}_n less than 0.1.

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To further identify the Pr and ET effects on SSM variability, we evaluate the relationships between their spectral slopes. Figure 5 shows the global distribution of SSM_{kw} , Pr_{kw} , and ET_{kw} expressed in terms of noise color in the three frequency bands based on SMAP and ERA5 data. We also evaluate the spectral slope of potential evaporation (PE_{kw}) from ERA5 to compare it with ET_{kw} .

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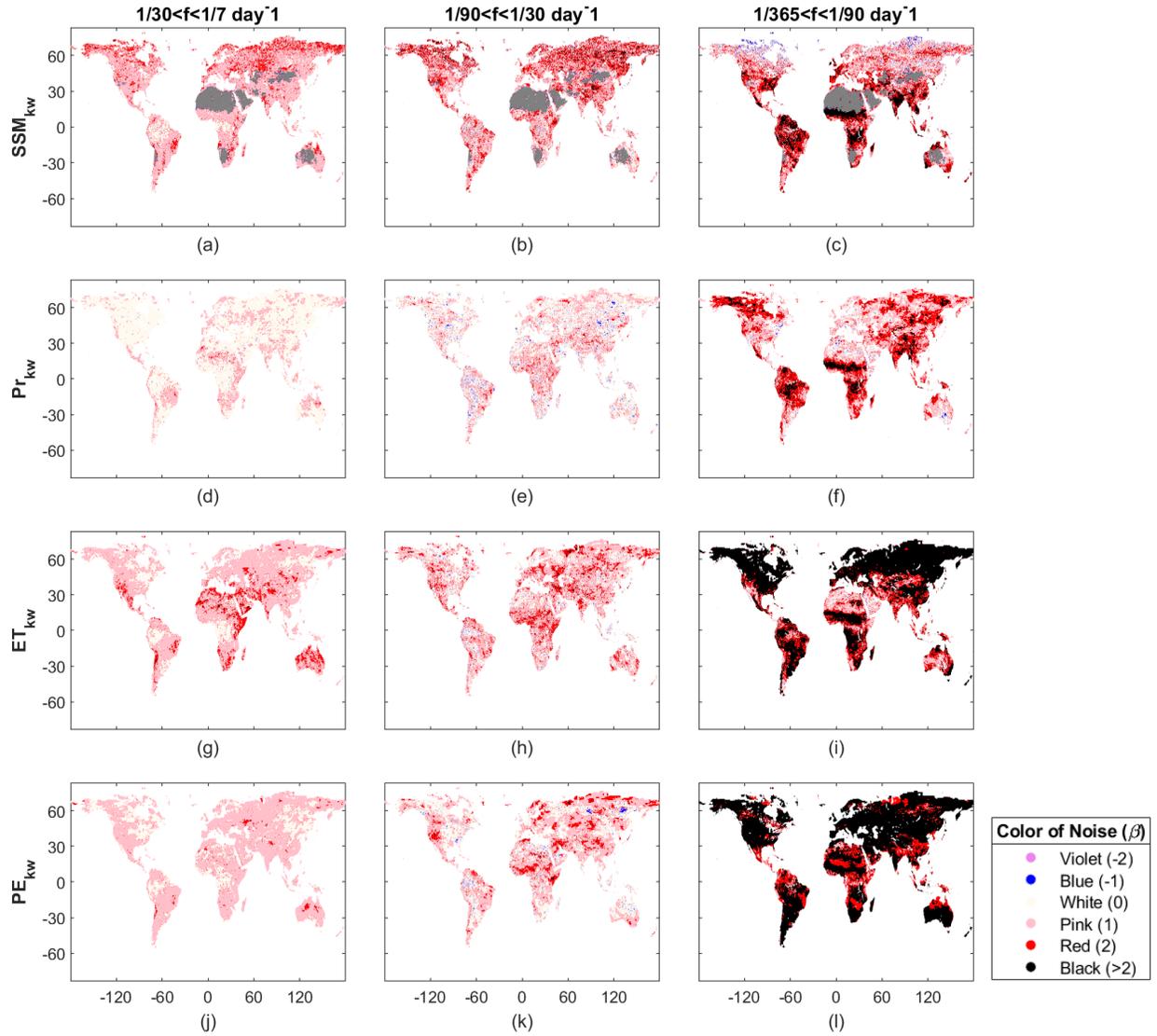
From a previous study (Xi et al., 2022), we have found that the low-frequency periodic components dominate the contribution to the variance of SSM, and it has more randomness on time scales shorter than monthly and more memory on time scales longer than seasonality. From Figure 5a-5f, we further find that there is a phase shift between SSM and Pr spectra in the two

440 higher frequency bands, especially the highest one, which implicates how Pr variability
441 propagates into the soil moisture system (Katul et al., 2007). In the weekly to monthly frequency
442 band where Pr is the dominant factor on SSM (according to Figure 4d), regions with smaller
443 Pr_{kw} lead to SSM spectra decay more rapidly. In most regions where Pr is similar to a white
444 noise, SSM exhibits a pink noise in the corresponding regions, indicating longer memory
445 induced by soil moisture (Salvucci and Entekhabi, 1994). In regions where Pr exhibits a pink
446 noise, like eastern Africa, Brazil, India, and northern Australia, SSM has a red noise spectrum
447 (Figures 5a and 5d). A similar relationship between SSM and Pr spectra can also be found in the
448 monthly to seasonal frequency band (Figures 5b and 5e), such as in southern North America,
449 southern and north-central Asia, and regions around the Mediterranean, but it is not as evident as
450 that in the highest frequency band since the effect of Pr on SSM variability decreases in this
451 frequency band (according to Figure 4e). In the seasonal to annual frequency band, ET performs
452 more effects on SSM variability than Pr for most regions (according to Figure 4f), so there are no
453 strong correlations between Pr and SSM spectra. In previous studies, soil moisture was found to
454 be similar to a red or black noise corresponding to precipitation having a white or pink noise at
455 high frequency (Katul et al., 2007; Nakai et al., 2014). The SSM_{kw} here is a little larger (Figure
456 5a). A possible reason is the effect of runoff. Since the only function of runoff is to prevent large
457 positive abnormalities in soil moisture, it may cause the time scale for soil moisture variability to
458 shorten (Delworth and Manabe, 1988). This mechanism will mainly affect the soil moisture
459 variability at high frequency and thus lead to less “redness” of soil moisture spectra.

460 Unlike between SSM_{kw} and Pr_{kw} , there is no such relationship between SSM_{kw} and
461 ET_{kw} , even at the highest frequency band where ET is dominant on SSM variability (Figures 5a-
462 5c and 5g-5i). It has been found that the sensitivity of soil moisture to precipitation and radiation

463 uncertainty performs differently in seasonality (Wei et al., 2008). Here we also find that Pr and
464 ET exert strong effects on SSM variability in different ways across different time scales. In
465 previous studies, unlike Pr serving as a forcing term, ET was shown to be related to the damping
466 term of soil moisture spectra (Delworth and Manabe, 1988; Katul et al., 2007; Nakai et al.,
467 2014), which modulates potential evaporation (PE). The differences between ET_{kw} and PE_{kw} are
468 mainly due to the variability of soil moisture. PE is an estimate of the maximum evaporation rate
469 from a surface of pure water for given meteorological conditions (Delworth and Manabe, 1988).
470 Weather fluctuations introduce a white or pink noise PE. However, unlike PE, ET is closely
471 related to soil moisture, emphasizing that soil moisture limits and regulates the supply of
472 moisture to the atmosphere on longer time scales. So the SSM dynamics influence ET spectra –
473 leading to a more red noise than PE spectra because SSM has a longer memory. This influence is
474 especially more visible in dry regions. The reason is that, compared to SSM in dry regions, SSM
475 in wet regions mostly tracks the variability of PE. So ET in wet regions will not be strongly
476 affected by SSM variability and thus still shows pink noise. On longer time scales, both ET and
477 PE show obvious seasonality that the low-frequency periodic components dominate the
478 contribution to the variance of signals (Figures 5i and 5l).

479 To summarize, the effects of Pr and ET on SSM variability are different across time
480 scales. In the two higher frequency bands (especially the weekly to monthly frequency band), Pr,
481 acting as a forcing by averaging the large oscillations to limit high-frequency components, has
482 more effects on SSM variability. In the seasonal to annual frequency band, ET, acting as the
483 dissipative process that prevents SSM anomalies from persisting indefinitely, has more effects on
484 SSM variability.



485

486 **Figure 5.** Noise color of SSM (Figure a-c), Pr (Figure d-f), ET (Figure g-i), and PE (Figure j-l) over the
 487 three frequency bands according to SSM_{kw} , Pr_{kw} , ET_{kw} , and PE_{kw} based on SMAP and ERA5 data. The
 488 colors in each figure represent the corresponding color of noise, referring to the power spectra of SSM,
 489 Pr, ET, and PE. The legend shows the color referring to each noise, and the number in brackets is the
 490 inverse number of the spectral slope of power-law noise corresponding to each noise color.

491 3.2 Comparison between CMIP5 simulations and SMAP and ERA5 references

492 Figure 6a-6c shows the average differences for H_{SEP_n} and H_{EEP_n} of model simulations

493 within CMIP5 compared to SMAP and ERA5 data. A significance test is performed and depicted
494 using stippling. Here, the “+” stippling means the region passes a 100% significance test, and the
495 “.” stippling means the region passes a 75% significance test. Therefore, we only focus on the
496 regions with stippling. For most regions, the multimodel differences of H_{SEP_n} are negative in the
497 two higher frequency bands and they are positive in the lowest frequency band, which means that
498 the CMIP5 simulations of the total effect of ET and Pr on SSM variability are smaller on time
499 scales shorter than seasonal and are larger on time scales longer than seasonal, compared to
500 SMAP and ERA5 data (Figure 6a-6c). The average difference of H_{SEP_n} is largest in the monthly
501 to seasonal frequency band (-0.6792 and -0.4492 with 100% and 75% significance) and smallest
502 in the weekly to monthly frequency band (-0.3365 and -0.2871 with 100% and 75% significance)
503 (Table 1). For all three frequency bands, the average differences of H_{SEP_n} are larger in Central
504 and Northern North America, Central and Eastern Europe, and regions near the equator.

Significance Frequency band (day ⁻¹)	100% significance test			75% significance test		
	1/7 ~ 1/30	1/30 ~ 1/90	1/90 ~ 1/365	1/7 ~ 1/30	1/30 ~ 1/90	1/90 ~ 1/365
BCC-CSM1.1	-0.2755	-0.5409	0.5554	-0.2249	-0.2797	0.4720
BNU-ESM	-0.3221	-0.6096	0.4718	-0.2740	-0.4001	0.4002
CanESM2	-0.3565	-0.7277	0.5245	-0.3057	-0.4523	0.4311
CNRM-CM5	-0.3323	-0.8718	0.5380	-0.2808	-0.6494	0.4492
CSIRO-Mk3.6	-0.3695	-0.8166	0.4851	-0.3220	-0.5809	0.4208
GFDL-CM3	-0.3262	-0.6595	0.4189	-0.2785	-0.4436	0.3417
GFDL-ESM2G	-0.3243	-0.6493	0.4494	-0.2762	-0.4388	0.3693
GFDL-ESM2M	-0.3250	-0.6566	0.4615	-0.2767	-0.4423	0.3796
MIROC5	-0.3330	-0.5357	0.3897	-0.2846	-0.2995	0.3119
MIROC-ESM	-0.3374	-0.6061	0.4011	-0.2884	-0.3794	0.3257
MIROC-ESM- CHEM	-0.3383	-0.6064	0.4021	-0.2893	-0.3814	0.3271

MRI-CGCM3	-0.3812	-0.8297	0.5968	-0.3321	-0.5989	0.5227
MRI-ESM1	-0.3804	-0.8293	0.5971	-0.3315	-0.5974	0.5225
NorESM1-M	-0.3089	-0.5702	0.4245	-0.2547	-0.3453	0.3421
Average (\pm)	-0.3365 \pm	-0.6792 \pm	0.4797 \pm	-0.2871 \pm	-0.4492 \pm	0.4011 \pm
1 SD)	0.0274	0.1110	0.0692	0.0280	0.1121	0.0685
Observation	0.4734	1.1492	0.4161	0.4305	0.9550	0.4741

505 **Table 1.** Observational and multimodel differences of H_{SEP_n} within CMIP5. The observational H_{SEP_n}

506 here is the original value without normalization across the three frequency bands.

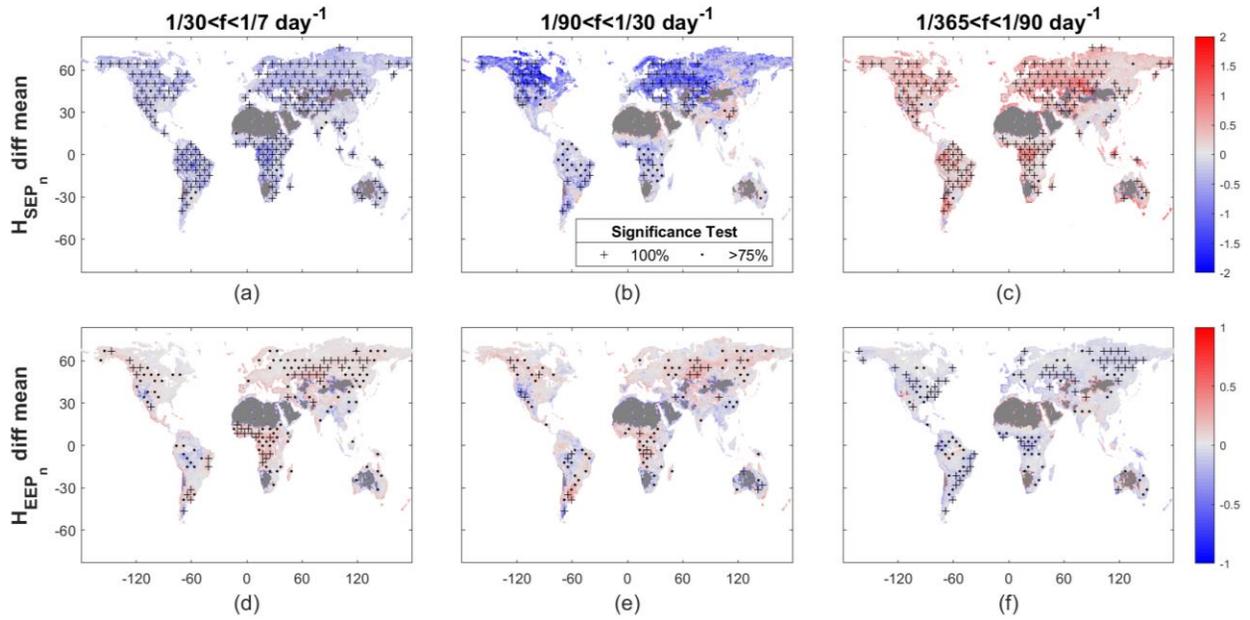
507 From section 3.1, we know that Pr dominates SSM variability in the two higher
508 frequency bands, and ET dominates it in the seasonal to annual frequency band. From Figure 6e-
509 6f, we find that in each frequency band, the effect of the corresponding dominant factor (i.e., Pr
510 or ET) on SSM simulated within the CMIP5 models tends to be smaller than that from ERA5
511 data. Specifically, in the two higher frequency bands where Pr is the dominant factor, models
512 overestimate the proportion of ET variability to the total variability of ET and Pr. Thus, the effect
513 of Pr on SSM is underestimated by models (Figure 6d-6e). In the lowest frequency band where
514 ET is the dominant factor, models underestimate the effects of ET on SSM. Unlike H_{SEP_n} , the
515 multimodel difference of H_{EEP_n} is largest in the weekly to monthly frequency band (-0.1259 and
516 -0.0770 with 100% and 75% significance) and smallest in the monthly to seasonal frequency
517 band (-0.0677 and -0.0515 with 100% and 75% significance) (Table 2). From Figure 6 (also
518 Table 1 and 2), CMIP5 simulations show larger differences on H_{SEP_n} than H_{EEP_n} , which means
519 that these CMIP5 models perform relatively well in capturing the proportion of ET and Pr
520 variability to their total variability, while they exhibit larger differences in simulating the total
521 effect of ET and Pr on SSM variability compared to SMAP and ERA5 data.

Significance	100% significance test	75% significance test
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Frequency band (day ⁻¹)	1/7 ~ 1/30	1/30 ~ 1/90	1/90 ~ 1/365	1/7 ~ 1/30	1/30 ~ 1/90	1/90 ~ 1/365
BCC-CSM1.1	0.1826	0.1089	-0.1059	0.1374	0.0992	-0.0579
BNU-ESM	0.1818	0.0927	-0.1414	0.1303	0.0751	-0.1141
CanESM2	0.1119	0.1245	-0.0800	0.0675	0.1146	-0.0524
CNRM-CM5	0.0891	0.0235	-0.0732	0.0493	0.0089	-0.0402
CSIRO-Mk3.6	0.0863	0.0575	-0.0793	0.0372	0.0386	-0.0517
GFDL-CM3	0.1216	0.0656	-0.0959	0.0738	0.0519	-0.0693
GFDL-ESM2G	0.1654	0.0988	-0.1087	0.1169	0.0866	-0.0833
GFDL-ESM2M	0.1637	0.0992	-0.1061	0.1133	0.0871	-0.0784
MIROC5	0.0778	0.0028	-0.0553	0.0159	-0.0272	-0.0249
MIROC-ESM	0.1110	0.0429	-0.0743	0.0542	0.0218	-0.0523
MIROC-ESM- CHEM	0.1100	0.0418	-0.0750	0.0527	0.0215	-0.0524
MRI-CGCM3	0.1126	0.0700	-0.0633	0.0715	0.0542	-0.0407
MRI-ESM1	0.1109	0.0711	-0.0624	0.0699	0.0550	-0.0392
NorESM1-M	0.1386	0.0492	-0.1003	0.0880	0.0336	-0.0792
Average (\pm 1 SD)	0.1259 \pm 0.0336	0.0677 \pm 0.0332	-0.0872 \pm 0.0227	0.0770 \pm 0.0347	0.0515 \pm 0.0374	-0.0597 \pm 0.0222
Observation	0.2406	0.4182	0.7747	0.2236	0.3756	0.7618

522 **Table 2.** Observational and multimodel differences of H_{EEP_n} within CMIP5. The observational H_{EEP_n}

523 here is the original value without normalization across the three frequency bands.



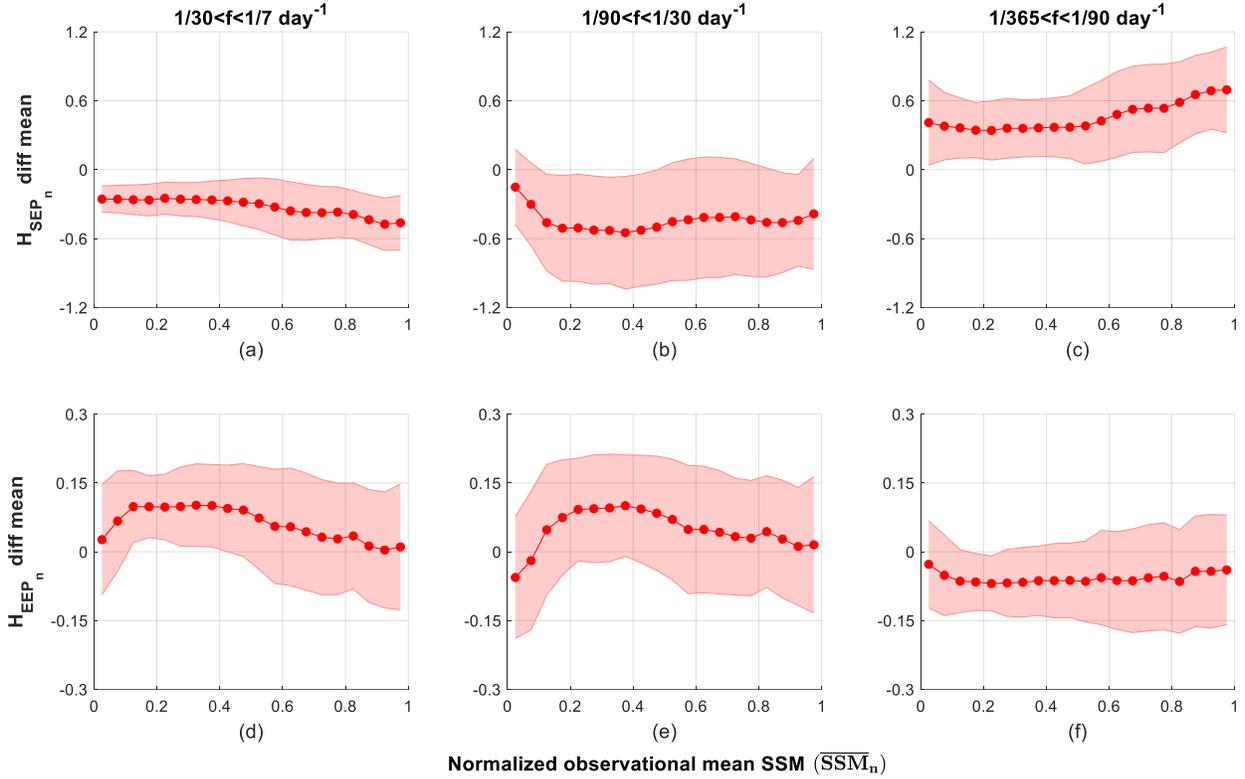
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525 **Figure 6.** Average differences of H_{SEP_n} (Figure a-c) and H_{EEP_n} (Figure d-f) between CMIP5 models and
 526 the observation-based data in the three frequency bands. Dark grey parts are regions with \overline{SSM}_n less than
 527 0.1. For each figure, “+” and “.” stippling represents the region that passes a 100% significance test and a
 528 75% significance test, respectively.

529 In addition to multimodel differences compared to SMAP and ERA5 data, the coefficient
 530 of variation (CV) of H_{SEP_n} and H_{EEP_n} across models are also investigated to estimate their
 531 statistical variance (Figure S4). For both H_{SEP_n} and H_{EEP_n} , the intermodel spread is larger in the
 532 weekly to monthly and monthly to seasonal frequency band and smaller in the seasonal to annual
 533 frequency band (also see Table S6). Therefore, for CMIP5 estimations of Pr and ET effect on
 534 SSM variability, there is a more extensive intermodel spread on time scales shorter than seasonal
 535 and a lower variance among models on time scales longer than seasonal time scale, suggesting an
 536 individual deficiency in representing the short-term variability and a systematic deficiency of
 537 these CMIP5 models in representing the long-term variability.

538 The multimodel differences of H_{SEP_n} and H_{EEP_n} are further analyzed with the mean SSM

539 on a global scale. To make a trade-off between high significance and the size of samplings, we
540 use the differences that pass a 75% significance test. Figure S3 shows the global distribution of
541 the mean SMAP SSM after spatiotemporal normalization (\overline{SSM}_n). For H_{SEP_n} (Figure 7a-7c),
542 models achieve their best estimates in transitional zones between dry and wet climates, where
543 there is both a strong coupling between soil moisture and Pr (Koster et al., 2004) as well as
544 between soil moisture and ET (Seneviratne et al., 2010). No matter whether \overline{SSM}_n increases or
545 decreases from the intermediate transitional zones, the differences of H_{SEP_n} increase. Therefore,
546 when considering the Pr and ET effect on SSM variability, the CMIP5 models can perform better
547 in regions with strong coupling between these variables, and the differences compared to
548 observation-based data tend to be more apparent in wet and dry regions where interactions are
549 weaker. This finding is particularly evident in the highest and lowest frequency bands where
550 observation-based H_{SEP_n} is smaller. On the other hand, from Figure 7d-7f, H_{EEP_n} differences
551 basically increase with the decrease of \overline{SSM}_n except for extremely dry regions, indicating that the
552 CMIP5 models have difficulties in estimating the interaction between Pr and ET in regions with
553 less soil moisture. When soil moisture is limited, ET is also limited, although sensitive to SSM.
554 Under this condition, ET variation is too small to impact climate variability, and the impact of Pr
555 variation on climate variability is almost independent on SSM as drier soils will lead to lower
556 precipitation likelihood (Seneviratne et al., 2010). Therefore, it is hard for models to capture
557 correct interactions between Pr and ET, shown as larger differences of H_{EEP_n} in drier regions. In
558 regions where \overline{SSM}_n is extremely low (less than 10%), models tend to correctly capture the
559 proportion of Pr and ET variability.



560

561 **Figure 7.** Comparison of average differences of H_{SEP_n} (Figure a-c) and H_{EEP_n} (Figure d-f) between
 562 CMIP5 models and observation-based data with \overline{SSM}_n in the three frequency bands. The red shading
 563 represents \pm one standard deviation. \overline{SSM}_n is separated into 20 bins of equal size (i.e., 0.05 for each
 564 bin), then the mean of H_{SEP_n} and H_{EEP_n} differences located in each bin (corresponding to the range of
 565 \overline{SSM}_n) were separately calculated for each frequency band. Differences in this figure are the values
 566 passing a 75% significance test. All values in the regions with \overline{SSM}_n less than 0.1 are removed.

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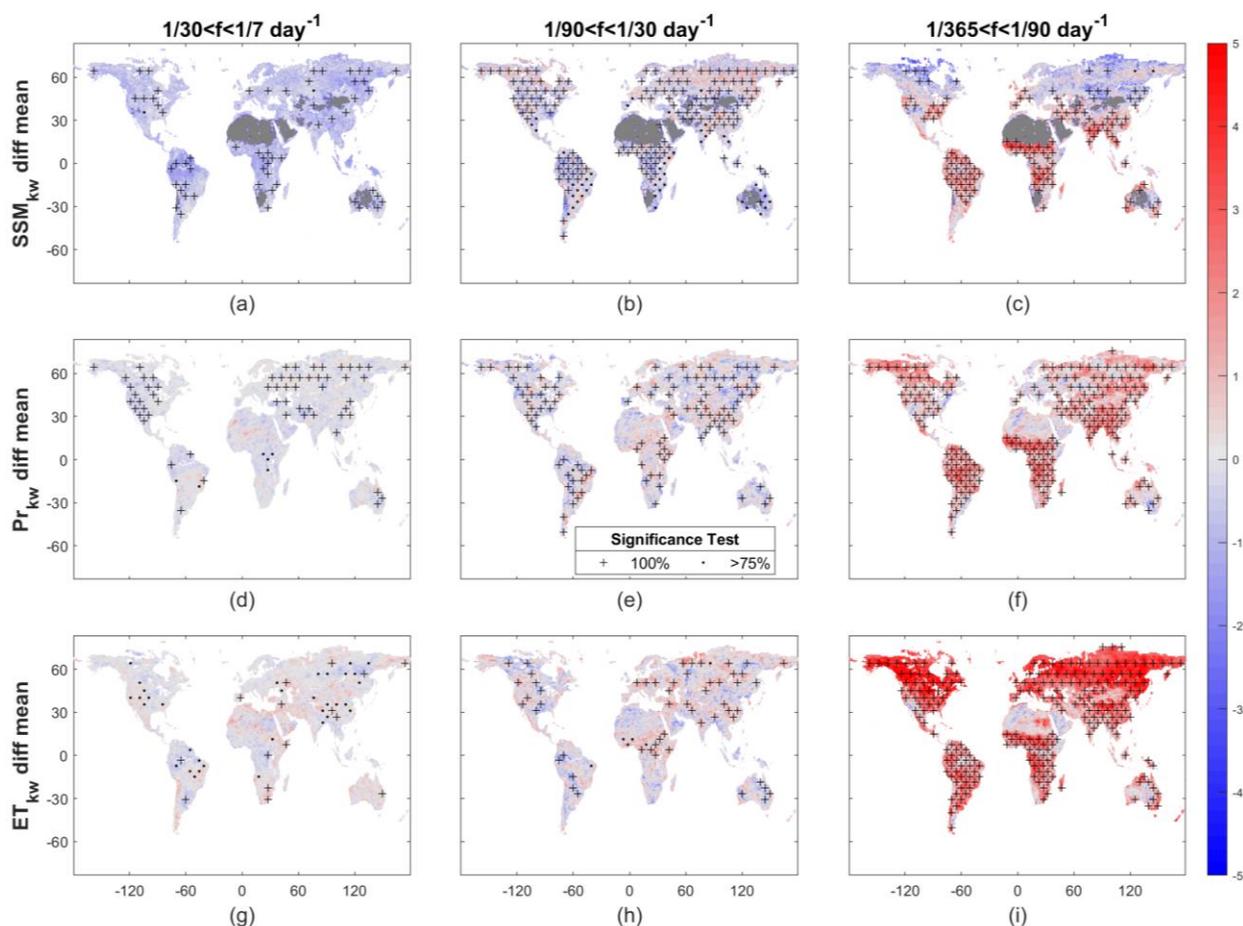
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Finally, we evaluate multimodel differences of the spectral slopes ($SSM_{k\omega}$, $Pr_{k\omega}$, and $ET_{k\omega}$) compared to the SMAP and ERA5 data (Figure 8). Negative differences mean that modeled spectra decay more rapidly and vice versa. Compared to SMAP SSM spectra, CMIP5 SSM spectra decay more rapidly in the two higher frequency bands and less rapidly in the seasonal to annual frequency band in most regions (Figure 8a-8c), indicating that these models underestimate the short-term variability and overestimate the long-term variability of SSM in a

573 non-linear way. For $Pr_{k\omega}$ (Figure 8d-8f) and $ET_{k\omega}$ (Figure 8g-8i), positive differences with high
 574 significance in most regions indicate that CMIP5 models underestimate their memory, implying
 575 land surface models may not be able to reproduce the correct intensity of Pr and ET variability,
 576 especially on time scales longer than seasonal. Our findings are aligned with previous studies
 577 (Katul et al., 2007; McColl et al., 2019; Nakai et al., 2014) but with different methods and
 578 models. We also find the differences characterizing the memory are not the same across
 579 frequencies and are most prominent in the seasonal to annual frequency band. This again
 580 suggests that models exhibit deficiency in representing long-term transpiration and soil moisture
 581 dynamics.



582

583 **Figure 8.** Average differences of $SSM_{k\omega}$ (Figure a-c), $Pr_{k\omega}$ (Figure d-f), and $ET_{k\omega}$ (Figure g-i) between
584 CMIP5 models and the observation-based data in the three frequency bands. Dark grey parts in Figure a-c
585 are regions with \overline{SSM}_n less than 0.1. For each subfigure, “+” and “.” stippling represents the region that
586 passes a 100% significance test and a 75% significance test, respectively.

587 3.3 Uncertainty analysis

588 Two parts during the data processing could introduce uncertainties to our analysis in this
589 study. First, since the SMAP data is non-continuous on the daily time scale, we fill the missing
590 values before performing Fourier analysis. The gap-filling process is the same as our previous
591 analysis and has been carefully validated using in-situ soil moisture data from International Soil
592 Moisture Network (Xi et al., 2022). Second, the interpolation during intermodel computation
593 may induce uncertainties since the spatial resolution of all these CMIP5 models are much coarser
594 than the “standard” spatial resolution (36 km×36 km, see section 2.7). Apart from comparing the
595 re-gridded results with an intermediate resolution (1°×1°) and finding that the differences are
596 very small (Xi et al., 2022), we also conduct a significance test to constrain the potential
597 uncertainties as much as possible. All statistical correlation analyses in this study are based on
598 the multimodel differences passing a high significance test (more than 75%), ensuring that a
599 systematic performance in land surface models is shown.

600 Apart from these technical issues, some other aspects may also cause uncertainties. One
601 issue is related to the potential biases of the SMAP data. This study conducts comparative
602 evaluations of ESMs within CMIP5 and uses SMAP products as the observations of SSM.
603 However, even though SMAP meets its performance target and has better performance than other
604 satellite products, its retrievals have been shown to exhibit potential errors in heavily vegetated
605 areas such as forests, with the presence of water bodies, and in frozen soil such as in the Arctic

606 tundra environment (Entekhabi et al., 2014; McColl et al., 2017; Wrona et al., 2017). Therefore,
607 when performing comparative assessments with model simulations, the biases in SMAP data
608 themselves should also be taken into consideration.

609 Another issue is the linear and time-invariant assumption of the interactions among SSM,
610 Pr, and ET. In this study, we assumed an LTI system of SSM, Pr, and ET and then performed the
611 Fourier analysis based on it. However, the relationships among them may not be linear and time-
612 invariant. For example, in regions with plenty of vegetation, precipitation is first intercepted by
613 the canopy, and then throughfall is further partitioned into surface runoff and infiltration water,
614 which directly affects SSM instead of precipitation. A previous study has also shown that there is
615 a higher linear relationship between soil moisture and precipitation in less-vegetated regions
616 (Sehler et al., 2019). Snow is another factor related to this issue. When the precipitation is snow,
617 it will not interact with SSM immediately. Instead, there is a snow accumulation and melting
618 process, which could take days, weeks, and even months. Thus, the relationship between SSM
619 and Pr may not be time-invariant in high-latitude regions.

620 Although we have mentioned that there are many complex physical processes involved in
621 the effects on SSM dynamics, and this study aims to only focus on the two elementary variables
622 related to SSM (i.e., Pr and ET), we still want to try to analyze the uncertainties from this aspect
623 and see how much confidence we could have under this LTI assumption. A feasible first step is to
624 mask the regions that could be most affected by these issues and see how the results will change.
625 This approach can also be used to quantify the uncertainties induced by SMAP data mentioned
626 above. We identify the regions with potential uncertainties as dense vegetation cover (vegetation
627 water content $> 5 \text{ kg/m}^2$), frozen landscapes (surface temperature $< 0^\circ\text{C}$), and the presence of
628 water bodies (water body fraction $> 5\%$ coverage of a pixel) (see Figure S5), which is similar to

629 a previous study (McColl et al., 2017). Then, we recalculate the observation-based H_{SEP_n} and
630 H_{EEP_n} , multimodel differences of H_{SEP_n} and H_{EEP_n} , and the CV of H_{SEP_n} and H_{EEP_n} across the
631 CMIP5 models with these regions being masked (Table S4, S5, and S6). We find that, although
632 being quantitatively inconsistent, these results are all qualitative across the three frequency bands,
633 illustrating the feasibility of our analysis on a global scale.

634 **4 Conclusions**

635 This study uses satellite-based observations to evaluate 14 Earth system models within
636 CMIP5 in simulating the effects of Pr and ET on SSM variability across three frequency bands.
637 We find that these models generally underestimate the total effects of Pr and ET on SSM in the
638 high-frequency bands (weekly to monthly and monthly to seasonal) and overestimate it in the
639 low-frequency band (seasonal to annual). Additionally, based on the findings that Pr dominates
640 weekly to seasonal SSM variability and ET dominates seasonal to annual SSM variability, these
641 models underestimate the effects on SSM by Pr or ET that is a dominant factor in each frequency
642 band. Across the three frequency bands, models perform better estimations in regions with strong
643 land-atmosphere interactions between the three variables. For the metrics investigated here,
644 models show an individual deficiency in representing short-term variability and a systematic
645 deficiency of long-term variability.

646 This study also identifies systematic metrics that can be used to assess model
647 performance and help refine process representation across time scales. Our results highlight that
648 the Earth system models within CMIP5 should improve their representation of precipitation and
649 evapotranspiration effects in modeling soil moisture.

650

651 **Appendix A:** Conceptual LTI systems representing the ET and Pr effects on SSM variability.

652 A transfer function (also known as system function) (Haykin and Van Veen, 2007)
653 mathematically represents the relationship between the input and output of a system (black-box
654 model). It can usually be used to describe the relationship between the signal excitation and
655 response of a linear time-invariant (LTI) system (Phillips et al., 2003) with the time-frequency as
656 a variable. For an LTI system, even if its specific structure and parameters are not known, its
657 model in the frequency domain can be regarded as a rational polynomial form. Then the
658 properties of the system can be determined by analyzing the input and output of the system.

659 LTI systems are subject to constraints of linearity and time invariance. The constraint of
660 linearity means that when multiple excitation signals act on the LTI system simultaneously, the
661 total response is equal to the sum of the corresponding individual effects of each excitation.
662 Besides, when the excitation increases by a specific multiple, the response also increases by the
663 same multiples. The constraint of time-invariance means that the response of the LTI system is
664 independent of the time the excitation acts on the system. This means that, regardless of the time
665 sequence of the input signal acting on the system, the output signals are the same. The only
666 difference is the time of their appearances. The constraint of linearity on the LTI system can be
667 expressed as:

$$T[ax_1(n) + bx_2(n)] = ay_1(n) + by_2(n) \quad (\text{A1})$$

668 where T represents the computational relationship of the system, $x_1(n)$, $y_1(n)$ and $x_2(n)$, $y_2(n)$
669 are two pairs of excitation and response, respectively. Besides, the constraint of time-invariance
670 on the LTI system can be expressed as:

$$y(n - m) = T[x(n - m)] \quad (\text{A2})$$

671 which means that when the excitation is delayed for a period of time m , the corresponding

672 response is also delayed for time m .

673 Subject to the constraints of linearity and time-invariance, if the signal applied to the LTI
674 system is decomposed (as an impulse signal), the response caused by the original excitation
675 signal is obtained by summing the responses generated by each component acting on the system.
676 In this way, the LTI system produces an output signal from any input signal, which can be
677 expressed as (considering the default system as a causal system):

$$y(t) = \int_0^{+\infty} x(\tau)h(t - \tau) = x(t) \otimes h(t) \quad (\text{A3})$$

678 where $x(t)$ and $y(t)$ are the input and output of the LTI system, respectively, and $h(t)$ is the
679 transfer function of the LTI system. According to the Convolution theorem, the convolution of
680 two signals in the time domain is equivalent to multiplying their corresponding spectra in the
681 frequency domain:

$$x(t) \otimes h(t) = X(k) \cdot H(k) \quad (\text{A4})$$

682 where $X(k)$ and $H(k)$ are the spectra of $x(t)$ and $h(t)$, respectively.

683 In the time domain, the terrestrial water balance can be simply expressed as:

$$\frac{dssm(t)}{dt} = pr(t) - et(t) - q(t) \quad (\text{A5})$$

684 where ssm is surface soil moisture, pr is precipitation, et is evapotranspiration, and q is
685 drainage and runoff. Neglecting drainage and runoff ($q = 0$), this water balance can be further
686 simplified as:

$$\frac{dssm(t)}{dt} = pr(t) - et(t) \quad (\text{A6})$$

687 where precipitation is the climate input to soil moisture, and evapotranspiration is the water
688 losses relative to soil moisture.

689 Any system that can be simulated as homogeneous linear differential equations with

690 constant coefficients can be regarded as an LTI system. In this way, the relationships between
691 SSM, ET, and Pr can be described assuming two conceptual LTI systems, where the inputs are
692 $et(t)$ and $pr(t)$, respectively, and the outputs are both $ssm(t)$ (Figure S2). Since the two
693 systems are both single-input and single-output (SISO) systems (Partington, 2004), we can focus
694 on the relationship between their excitations and responses without caring about the internal
695 variations of the systems. In this way, the relationships between excitation and response of the
696 two LTI systems can be expressed as:

$$ssm(t) = et(t) \otimes h_{se}(t) \quad (A7)$$

$$ssm(t) = pr(t) \otimes h_{sp}(t) \quad (A8)$$

697 where $h_{se}(t)$ and $h_{sp}(t)$ are the transfer function of the “ET-SSM” LTI system (Figure S2(a))
698 and “Pr-SSM” LTI system (Figure S2(b)), respectively.

699 It is hard to investigate these two transfer functions in the time domain. However, by
700 applying the convolution operator, equations (A4) and (A5) in the time domain can be converted
701 into the frequency domain as a product:

$$F_{SSM}(k) = F_{ET}(k) \cdot H_{SE}(k) \quad (A9)$$

$$F_{SSM}(k) = F_{Pr}(k) \cdot H_{SP}(k) \quad (A10)$$

702 where $H_{SE}(k)$ and $H_{SP}(k)$ are the Fourier transforms of the transfer functions $h_{se}(t)$ and $h_{sp}(t)$,
703 respectively. The two LTI systems change the spectra of the input signal by weighting each of its
704 frequency components. This change is completely determined by the transfer functions $H_{SE}(k)$
705 and $H_{SP}(k)$, which serve as a weighting function transforming the excitation with the spectrum
706 of $F_{ET}(k)$ and $F_{Pr}(k)$ into the response with the spectrum of $F_{SSM}(k)$.

707 Assuming the input of the two systems is a power signal (i.e., signal power is finite),
708 equations (A6) and (A7) can be read in terms of the power spectrum as:

$$E_{SSM}(k) = E_{ET}(k) \cdot |H_{SE}(k)|^2 \quad (\text{A11})$$

$$E_{SSM}(k) = E_{Pr}(k) \cdot |H_{SP}(k)|^2 \quad (\text{A12})$$

709 In this way, the effects of ET and Pr variability on SSM variability can be identified by
 710 $|H_{SE}(k)|^2$ and $|H_{SP}(k)|^2$, respectively.

711 To consider both ET and Pr effects on SSM, subject to linearity constraints, the two
 712 conceptual LTI systems can be combined as (Figure S2(c)):

$$ssm(t) = et(t) \otimes h_{se}(t) + pr(t) \otimes h_{sp}(t) \quad (\text{A13})$$

713 If we use an identical transfer function to replace the two cascaded transfer functions as the
 714 internal mechanism to capture the total effects of ET and Pr on SSM:

$$ssm(t) = et(t) \otimes h_{sep}(t) + pr(t) \otimes h_{sep}(t) \quad (\text{A14})$$

715 where $h_{sep}(t)$ is the transfer function of the LTI system shown in Figure 2 and can be performed
 716 by spectral analysis as the two SISO systems discussed above.

717

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729

730 **Data Availability Statement**

731 The codes and data for analysis in this study are available at

732 <https://purr.purdue.edu/publications/3999/1>.

733

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